# Lorentz Invariance Violation and Extra Dimensions

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In collaboration with M. Libanov, I. Gordeli

0712.1136 0901.4347 0904.0509, work in progress

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# Outline

#### Extra Dimensions as Alternative to Compactification

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Motivation Braneworlds with broken LI

#### Spectra of Perturbations

Localization and Delocalization Special Cases

 Does not have to be something fundamental (strings, branes, etc)

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$$\int dt d^d x \sqrt{g} \left( (\partial_0 \phi)^2 - (\Delta \phi)^2 - \epsilon (\partial_i \phi)^2 \right)$$

becomes Lorentz invariant in IR

#### Lorentz Invariant case



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Metric [PK Libanov]

$$ds^{2} = e^{-2k\xi|z|}dt^{2} - e^{-2k\zeta|z|}d\mathbf{x}^{2} - dz^{2}$$

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Matter - ideal relativistic fluid

$$T_B^A = u^A u_B(p+\rho) - p \delta_B^A + \Lambda, \quad u_A u^A = 1$$

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$$\rho = -\Lambda + 6k^2\zeta^2$$
$$w = -1 + \frac{3\zeta^2 - 2\zeta\xi - \xi^2}{\rho}$$
$$\omega = -1 + \frac{3\zeta(\zeta - \xi)}{\rho}$$

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But brane NEC implies  $\zeta - \xi < 0$ 

# **Microscopic Solution**

2+1 Lifshitz model [Kachru Liu Mulligan]

$$S = -\int rac{1}{e^2} F_{(2)} \wedge *F_{(2)} + F_{(3)} \wedge *F_{(3)} - c \int F_{(2)} \wedge B_{(2)} \,,$$

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Solution of Einstein equations  $Z = \zeta/\xi$ 

$$\Lambda = -\frac{Z^2 + Z + 4}{2L^2}$$
$$A^2 = \frac{2Z(Z - 1)}{L^2}$$
$$B^2 = \frac{4(Z - 1)}{L^2}$$

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To avoid tachyonic solutions Z > 1

# Correspondence

#### [PK Gordeli]

	KL	KLM
Cosmological constant	$\Lambda = -\rho - 3k^2$	$\Lambda = -L^{-2}(Z^2 + Z + 4)$
First component	W	$-1+rac{A^2+B^2}{2(-\Lambda-3L^{-2})}$
Second component	ω	$-1+rac{B^2}{2(-\Lambda-3L^{-2})}$
LIV parameter	$w - \omega$	A
Anisotropy	$p_1 - p_4$	Energy flux A <sup>2</sup>
		w.r.t z direction
Constraints	Reality of Fluxes	Null energy condition

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 $\ref{eq:linear}$  Recall that in these examples bulk NEC is inconsistent with brane NEC  $\ref{eq:linear}$ 

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The statement does not depend on the volume  $\int_{-\infty}^{+\infty} \sqrt{g} \, dz$  of the extra dimension

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$$a'' = b'' = egin{cases} (z - z_0)^2, & 0 \leq z < z_0, \ 0, & z \geq z_0. \end{cases}$$

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But one needs to have  $\kappa/k \sim 1$ 

In LI case everything is fine, e.g.

$$a'' = b'' = egin{cases} (z - z_0)^2, & 0 \leq z < z_0, \ 0, & z \geq z_0. \end{cases}$$

But still bulk matter needs to be localized near brane  $k \neq 0$  entails

$$w(z) = -\frac{(\xi^2 + 2\xi\zeta + 3\zeta^2)e^{-2\zeta kz} - 4\kappa^2/k^2}{6\zeta^2 e^{-2\zeta kz} - 12\kappa^2/k^2}$$
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But one needs to have  $\kappa/k \sim 1$  which makes the model useless phenomenologically

# Outlook so far

 Coordinate independent description of spaces with broken Lorentz Invariance (canonical formulation)

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- Coordinate independent description of spaces with broken Lorentz Invariance (canonical formulation)
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Obtain these solutions from String Theory

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- Obtain these solutions from String Theory
- So what is the dynamics in these backgrounds?

Spectra of Perturbations

Metric

$$ds^{2} = e^{-2a(z)}dt^{2} - e^{-2b(z)}d\mathbf{x}^{2} - dz^{2}$$

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Geodesic equation

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for fifth component

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Which means a' < b' – localization and a' > b' – delocalization

Scalar 5D field

$$S = \int dt d\mathbf{x} \int_{-\infty}^{+\infty} dz \sqrt{g} g^{AB} \partial_A \phi \partial_B \phi$$

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Fourier transform, redefinition of  $\phi,$  and reparametrization of z yield Schrödinger equation

$$\chi'' + (E^2 - V)\chi = 0$$

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potential

$$V = -\frac{1}{4}a'^{2} + \frac{9}{4}b'^{2} + p^{2}e^{2(b-a)} - \frac{1}{2}(a''+3b'')$$

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- The behavior at infinity which is controlled by the Lorentz invariance violation (the sign of b − a). The potential V can increase/decrease as y → ∞.
- The sign of the delta-function term. The potential may have either delta-well or delta-peak depending on this sign.

▶ If the momenta-dependent term increases as  $y \to \infty$ , then one has a discrete spectrum as in the box-type potential. The potential might have local minima and maxima but the behavior of this potential at infinity qualitatively defines the character of the spectrum. On the contrary, if the potential decays at infinity, then we have continuous spectrum of plane waves propagating along y-direction. Some combination of these two scenarios is possible when  $V \rightarrow V_{\infty} = const$  as  $z \to \infty$ . Then those modes with  $E^2 < V_{\infty}$  belong to discrete spectrum and modes with  $E^2 > V_{\infty}$  contribute to continuous spectrum.

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- The sign of delta-function term affects zero mode existence. In a delta-well there might be a zero-mode and none in a delta-peak.

# Linear Parameterization

$$a(z) = \xi k|z|, \ b(z) = \zeta k|z|$$

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# Linear Parameterization

$$a(z) = \xi k |z|, \ b(z) = \zeta k |z|$$
  
Potential

$$V_{p}(y) = \frac{p^{2} (1 + \xi k |y|)^{2(1 - \zeta/\xi)}}{4(1 + \xi k |y|)^{2}} - 3\zeta k \delta(y)$$

# Potentials



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Space of Metrics



Metric

$$ds^2 = dt^2 - e^{-2k|z|} d\mathbf{x}^2 - dz^2$$

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$$ds^2 = dt^2 - e^{-2k|z|} d\mathbf{x}^2 - dz^2$$

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$$V(z) = p^2 \mathrm{e}^{2k|z|} + \frac{9}{4}k^2 - 3k\delta(z)$$

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Generic solution

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Matching

$$\frac{p}{k}\frac{\mathsf{K}_{\nu+1}\left(\frac{p}{k}\right)}{\mathsf{K}_{\nu}\left(\frac{p}{k}\right)} = \frac{3}{2} + \nu$$

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# Spectrum of Model A



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Zero mode at small momenta

$$E^2 = 3p^2 \left(1 - \frac{p}{k} + \mathcal{O}(p^2)\right)$$

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higher modes at low momenta

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• large momenta – everything  $E = p + \dots$ 

### Static Potential

Brane-to-Bulk Propagator

$$\left[-\partial_z^2 - E^2 + p^2 e^{2k|z|} + \frac{9}{4}k^2 - 3k\delta(z)\right]\Delta_p(E, p, z) = \delta(z)$$

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Put E = 0 to find the Green function. On the brane z = 0

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Brane static potential

$$G(r)=\frac{k}{4\pi r}\left(1+\frac{2}{\pi kr}\right)$$

# Fermions

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Action

$$S=\int dt d{f x} \int\limits_{-\infty}^{+\infty} dz\,\sqrt{g}\left(iar{\Psi}ar{
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Dirac equation

$$i\Gamma^A 
abla_A \Psi(x,z) + m_\psi \Psi(x,z) = 0, \quad m_\psi = m \operatorname{sign}(z)$$
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In rescaled variables

$$\left(E\gamma^{0}-\mathrm{e}^{k|z|}\gamma^{i}p_{i}-\gamma^{5}\partial_{z}\right)\psi-m_{\psi}\psi=0$$

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Using special rep os  $\gamma\text{-matrices}$ 

$$E\chi - ip\sigma_3\phi e^{k|z|} + \sigma_1\phi' - m_\psi\phi = 0$$
  
$$E\phi + ip\sigma_3\chi e^{k|z|} - \sigma_1\chi' - m_\psi\chi = 0$$

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## **Dispersion** Relations

• At m > k/2 the dispersion relation is

$$E\simeq\left(rac{2m}{2m-k}
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► At 0 < m < k/2</p>

$$E \simeq 2m \frac{\Gamma\left(\frac{1}{2} - \frac{m}{k}\right)}{\Gamma\left(\frac{1}{2} + \frac{m}{k}\right)} \left(\frac{p}{2k}\right)^{\frac{2m}{k}}$$

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#### Higher modes

Higher modes at low momenta

$$E_n = m \sqrt{1 + \left(\frac{\pi nk}{k - 2\Psi\left(\frac{1}{2}\right)m + 2m\log\frac{p}{2k}}\right)^2}$$

The case m = 0 is very neat

$$E_n = -rac{\pi k(2n+1)}{4\log rac{p}{2k}} + \mathcal{O}(p), \quad n \in \mathbb{N}$$

All modes at high momenta

$$E = p + \mathcal{O}(p^{1/3})$$

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Left and Right fermions  $\psi_{L,R} = \frac{1}{2}(1 \mp \gamma_5)\psi$ 

$$\psi^{>}_{\boldsymbol{L},\boldsymbol{R}} = \left(rac{\xi_+(t\mathrm{e}^{k|z|})}{\xi_+(t)} \pm \gamma rac{\xi_-(t\mathrm{e}^{k|z|})}{\xi_-(t)}
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Both left-handed and right-handed components do not vanish as functions of *z*.

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### Outlook

Study other models with broken LI

Holography

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- Holography
- Gravity and Cosmology

## Outlook

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- Holography
- Gravity and Cosmology
- Low dimensional SUSY
- ▶ ...