

The Universe Black Holes & Entanglement

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$$|\uparrow\rangle|\downarrow\rangle + |\downarrow\rangle|\uparrow\rangle$$

The Universe

<https://scaleofuniverse.com>

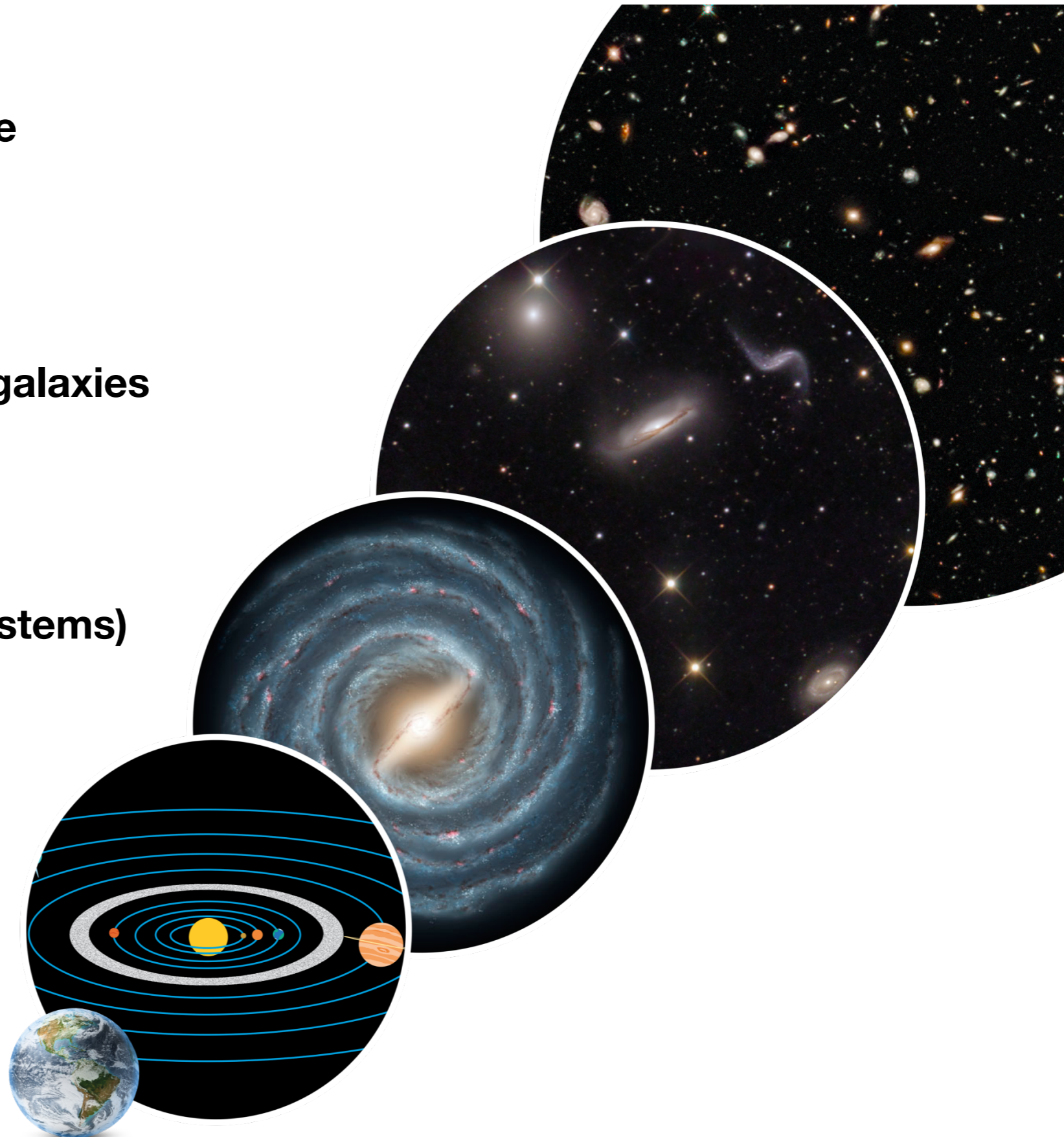
Large Structures of the Universe

Clusters (and superclusters) of galaxies

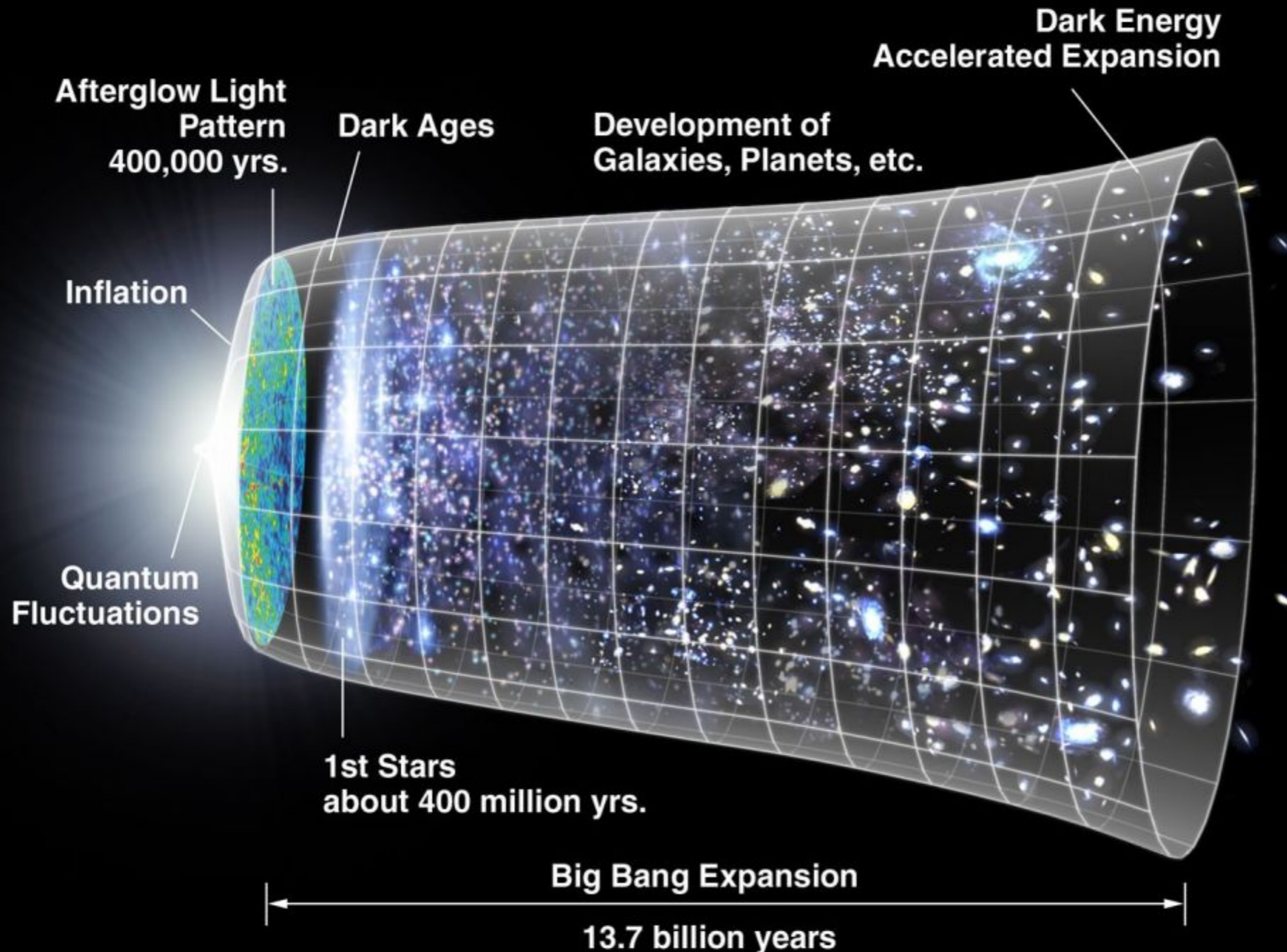
Galaxies (100 billions of solar systems)

Solar systems

Planets



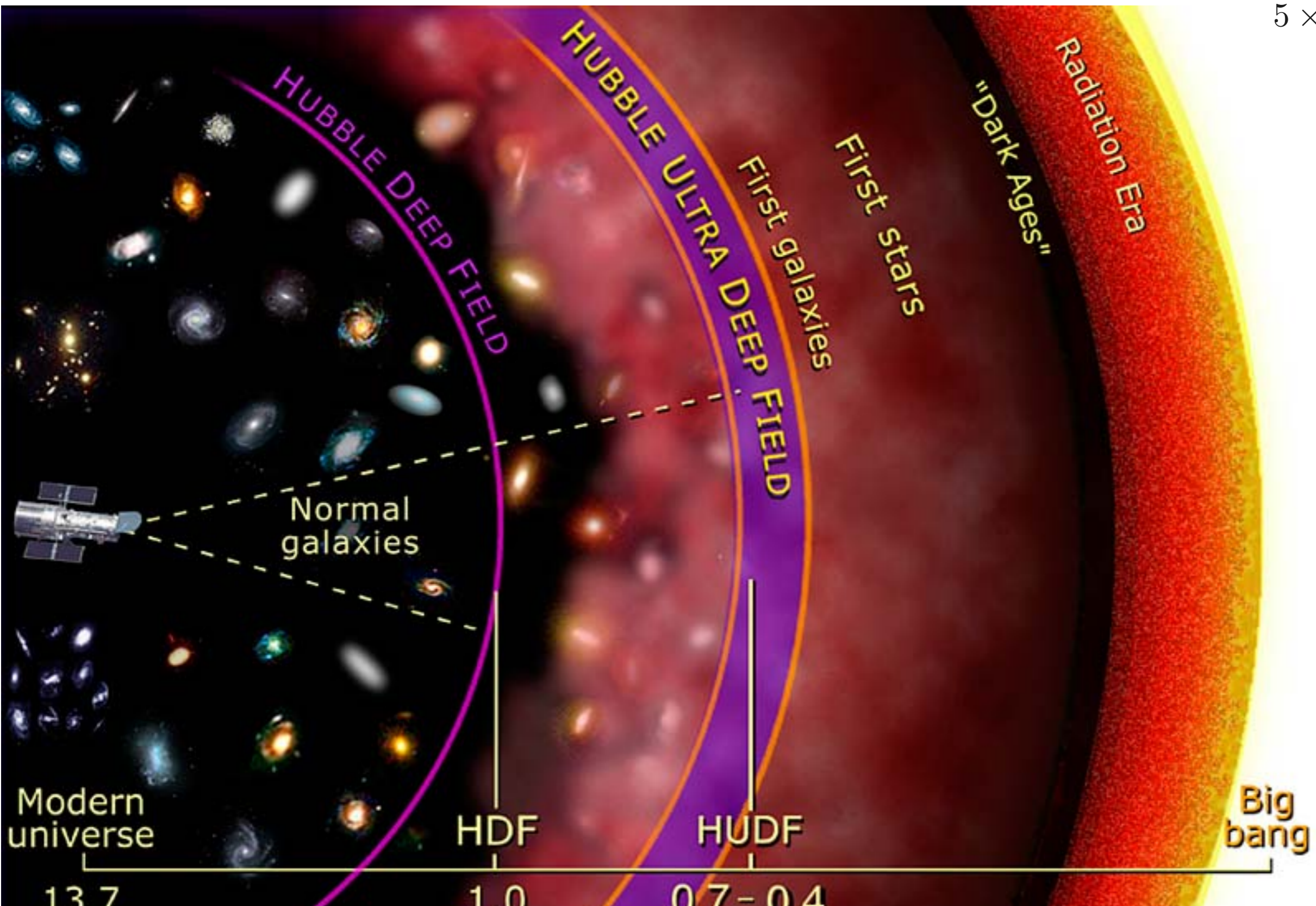
Expanding Universe



How far we can see

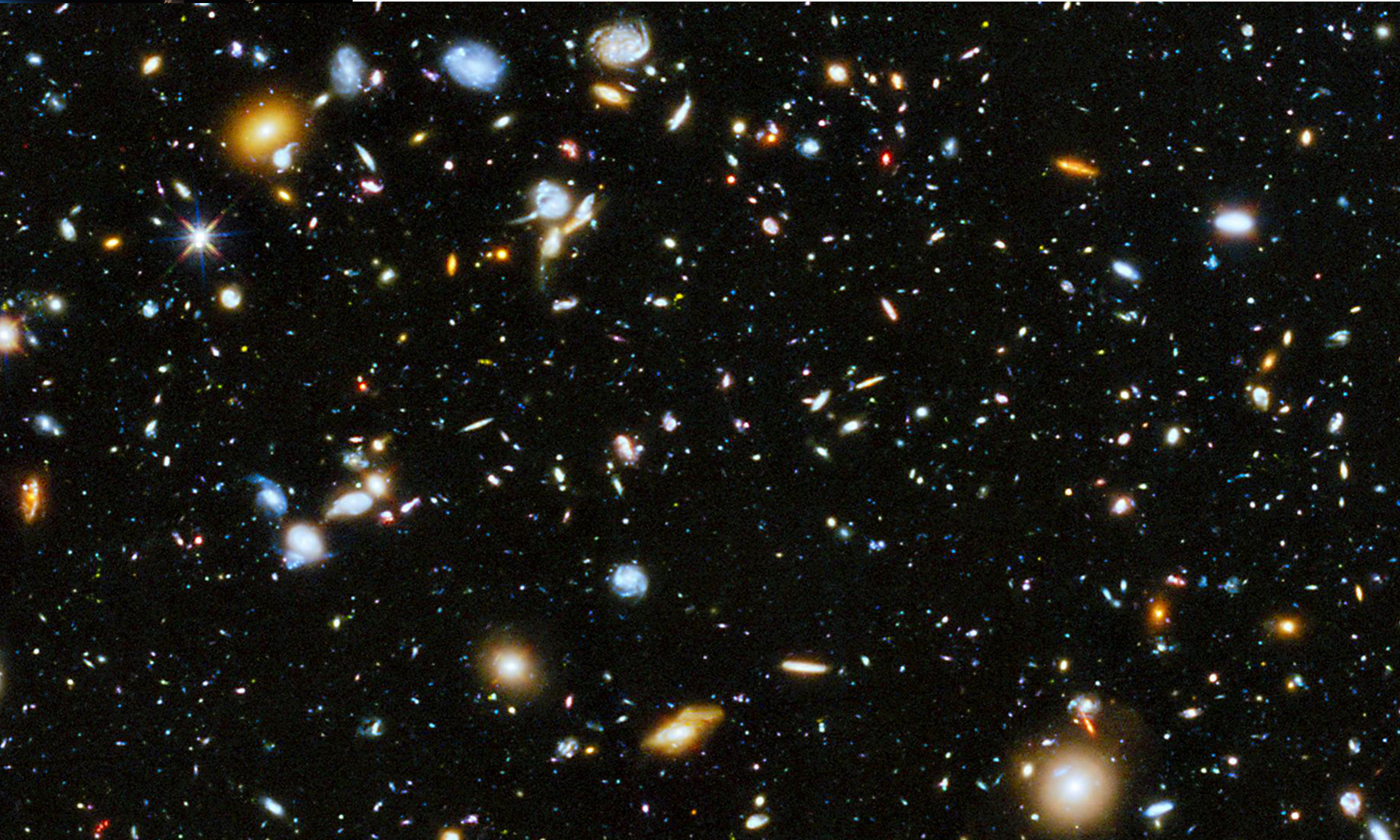
~13.2 billion years ago

$$5 \times 10^{26} m$$

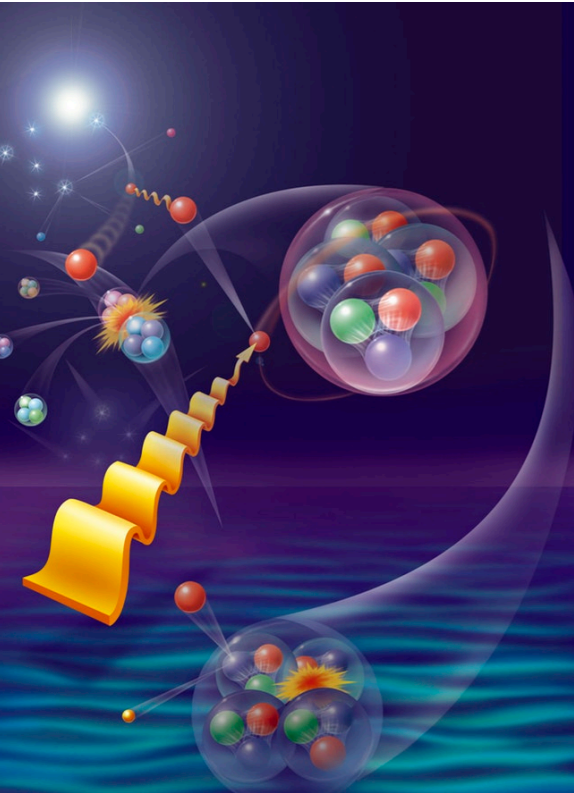




Hubble Ultra Deep Field

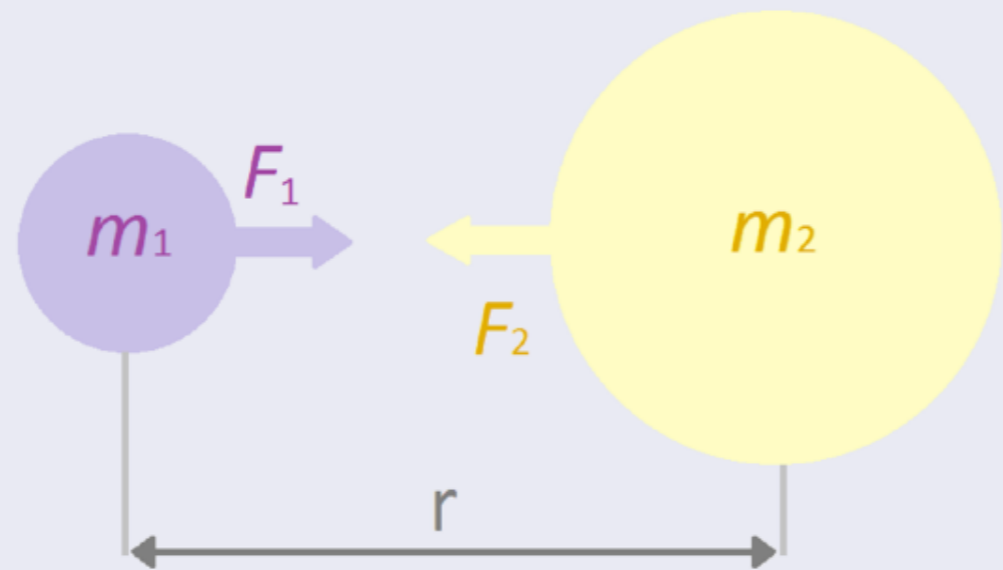


Scales in Physics

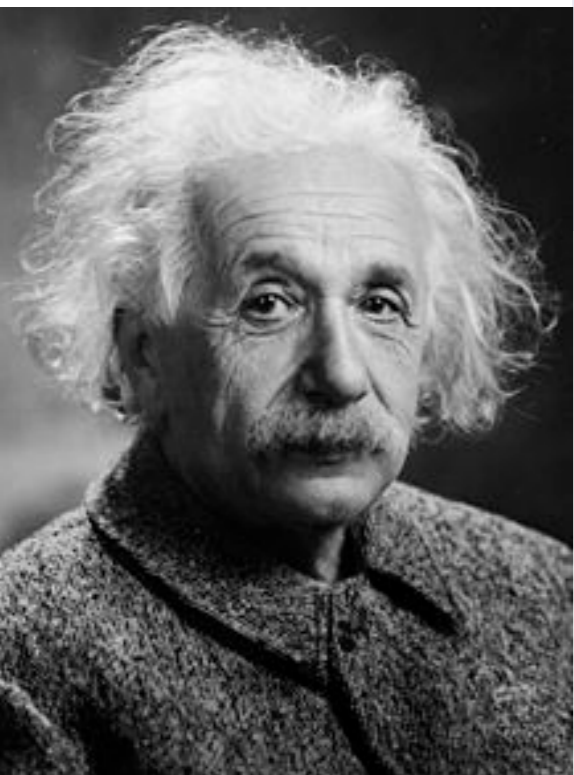
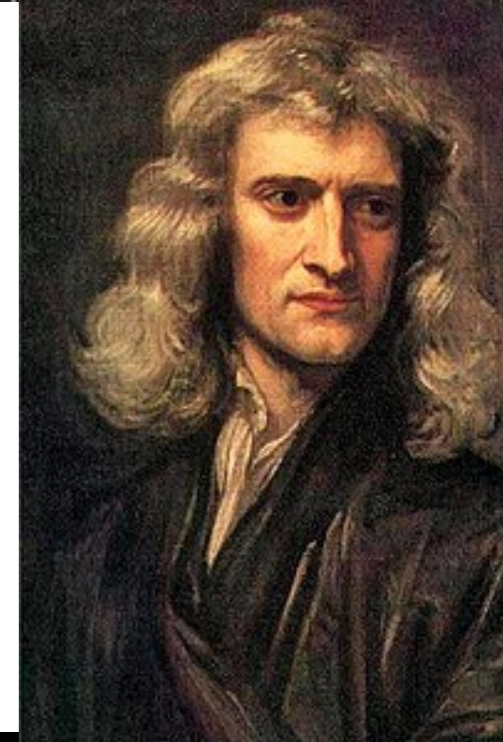


$$\leq 10^{-5}m$$

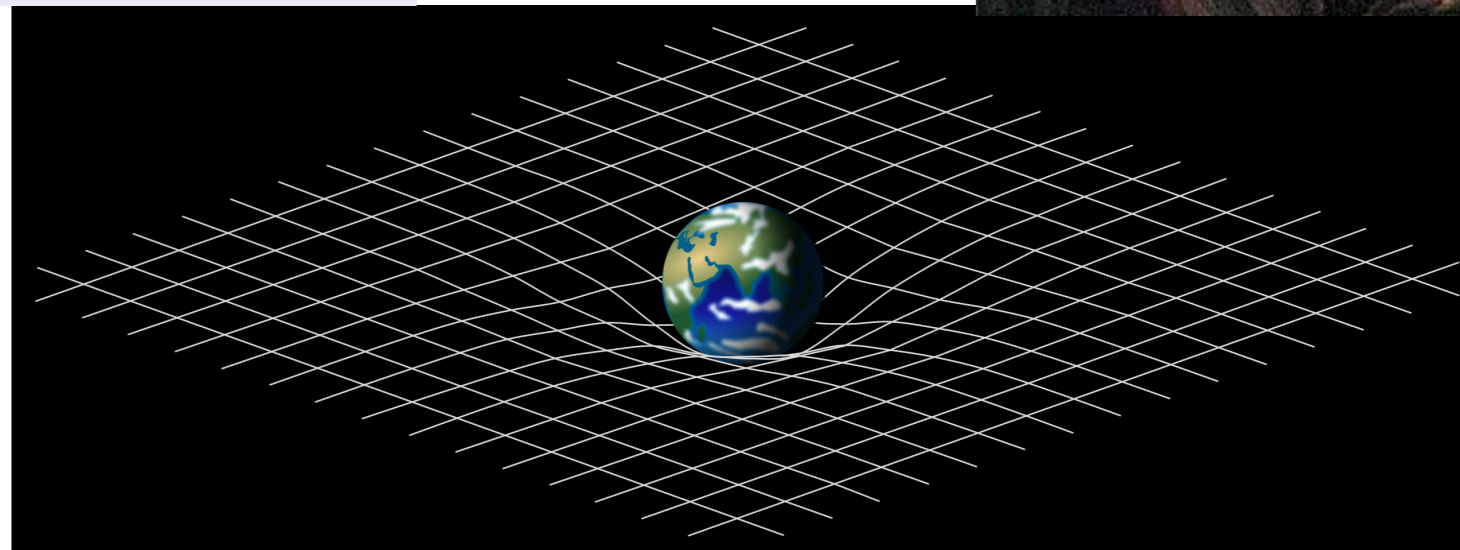
Quantum Mechanics
Quantum Field Theory



$10^{-5} - 10^{20}m$
Newton theory

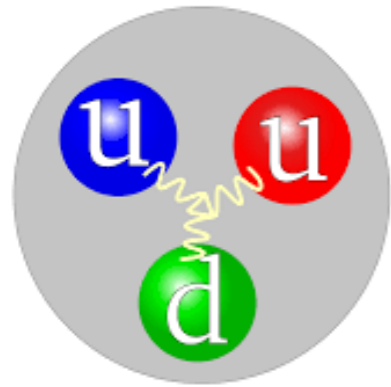


$\geq 10^{20}m$
General Relativity

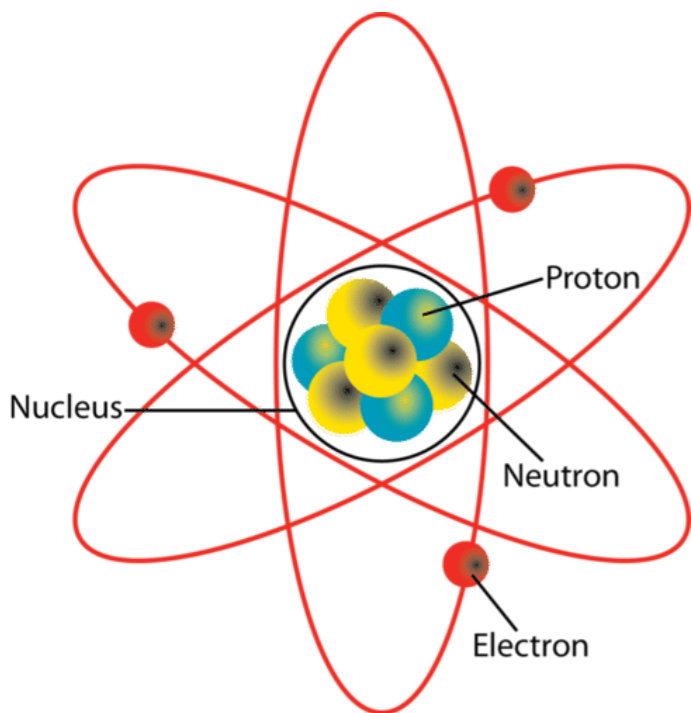


The Standard Model

proton

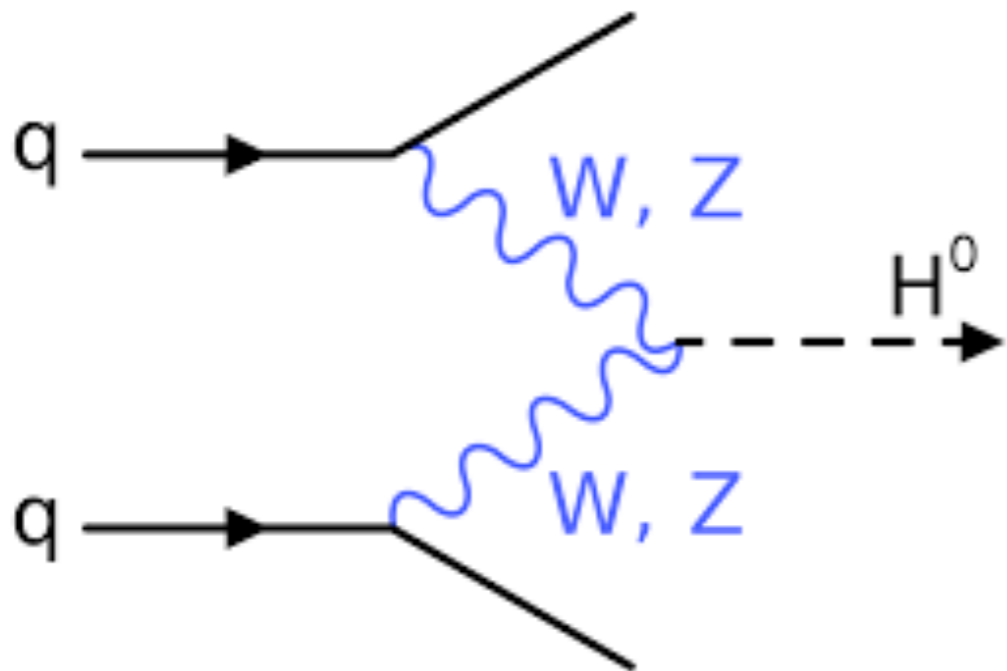
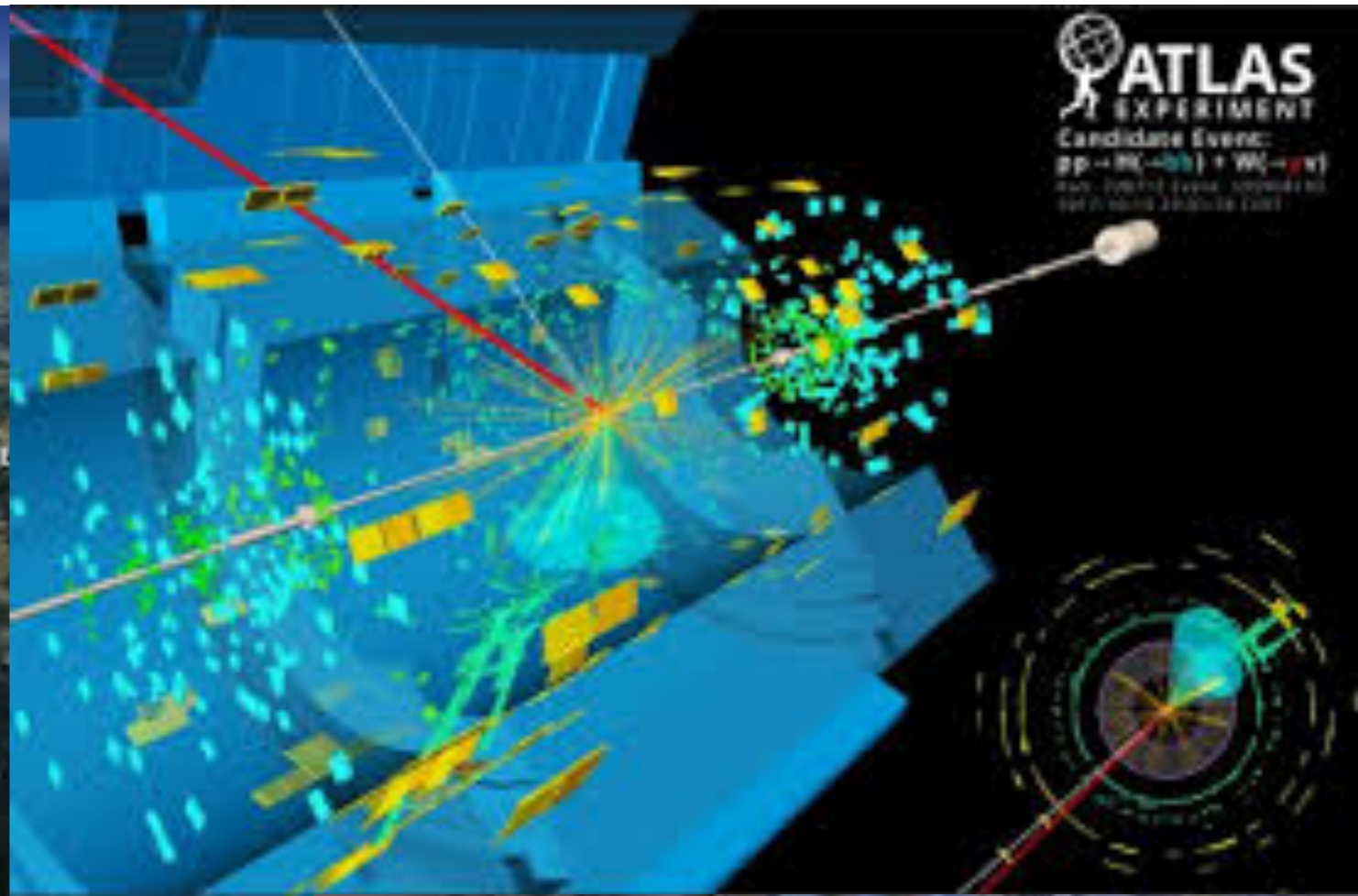


atom



	mass → $\approx 2.3 \text{ MeV}/c^2$ charge → $2/3$ spin → $1/2$	mass → $\approx 1.275 \text{ GeV}/c^2$ charge → $2/3$ spin → $1/2$	mass → $\approx 173.07 \text{ GeV}/c^2$ charge → $2/3$ spin → $1/2$	mass → 0 charge → 0 spin → 1	mass → $\approx 126 \text{ GeV}/c^2$ charge → 0 spin → 0
	u up	c charm	t top	g gluon	H Higgs boson
QUARKS	mass → $\approx 4.8 \text{ MeV}/c^2$ charge → $-1/3$ spin → $1/2$	mass → $\approx 95 \text{ MeV}/c^2$ charge → $-1/3$ spin → $1/2$	mass → $\approx 4.18 \text{ GeV}/c^2$ charge → $-1/3$ spin → $1/2$	mass → 0 charge → 0 spin → 1	
	d down	s strange	b bottom	γ photon	
LEPTONS	mass → $0.511 \text{ MeV}/c^2$ charge → -1 spin → $1/2$	mass → $105.7 \text{ MeV}/c^2$ charge → -1 spin → $1/2$	mass → $1.777 \text{ GeV}/c^2$ charge → -1 spin → $1/2$	mass → $91.2 \text{ GeV}/c^2$ charge → 0 spin → 1	
	e electron	μ muon	τ tau	Z Z boson	
	mass → $< 2.2 \text{ eV}/c^2$ charge → 0 spin → $1/2$	mass → $< 0.17 \text{ MeV}/c^2$ charge → 0 spin → $1/2$	mass → $< 15.5 \text{ MeV}/c^2$ charge → 0 spin → $1/2$	mass → $80.4 \text{ GeV}/c^2$ charge → ± 1 spin → 1	
	ν_e electron neutrino	ν_μ muon neutrino	ν_τ tau neutrino	W W boson	
					GAUGE BOSONS

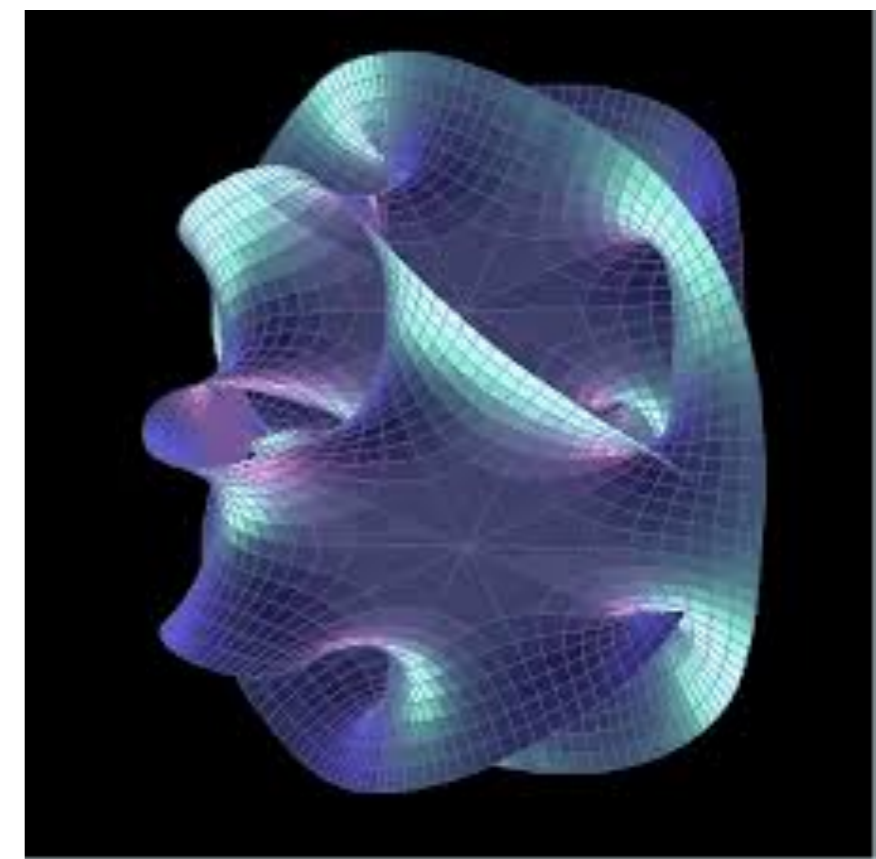
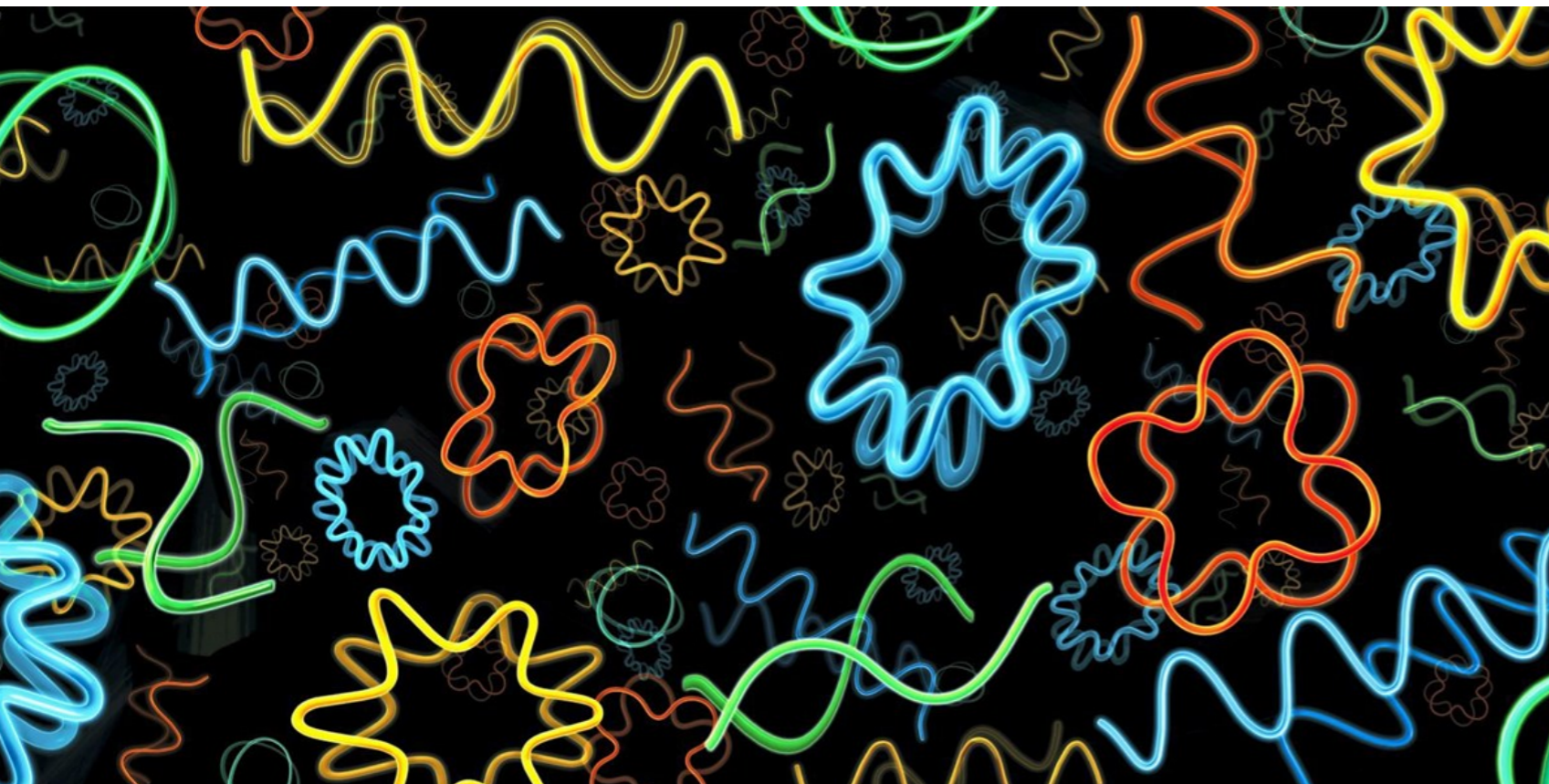
LHC



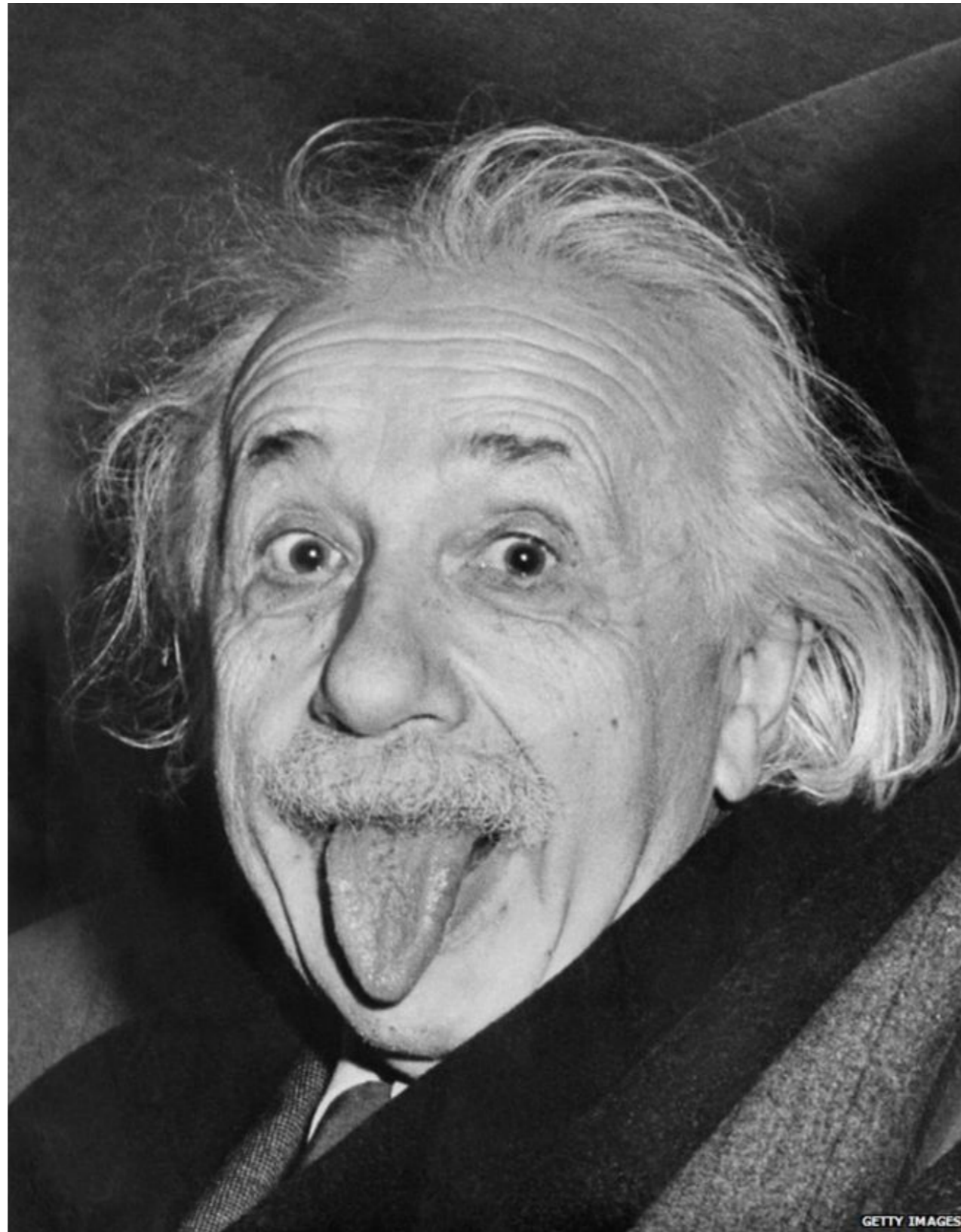
One of the fundamental problems of modern theoretical physics is to connect *quantum mechanics* with *general relativity*

Best candidate — **String Theory**

Describes interaction of strings (1+1 dimensional objects) traveling inside 9+1 dimensional spacetime



General Relativity



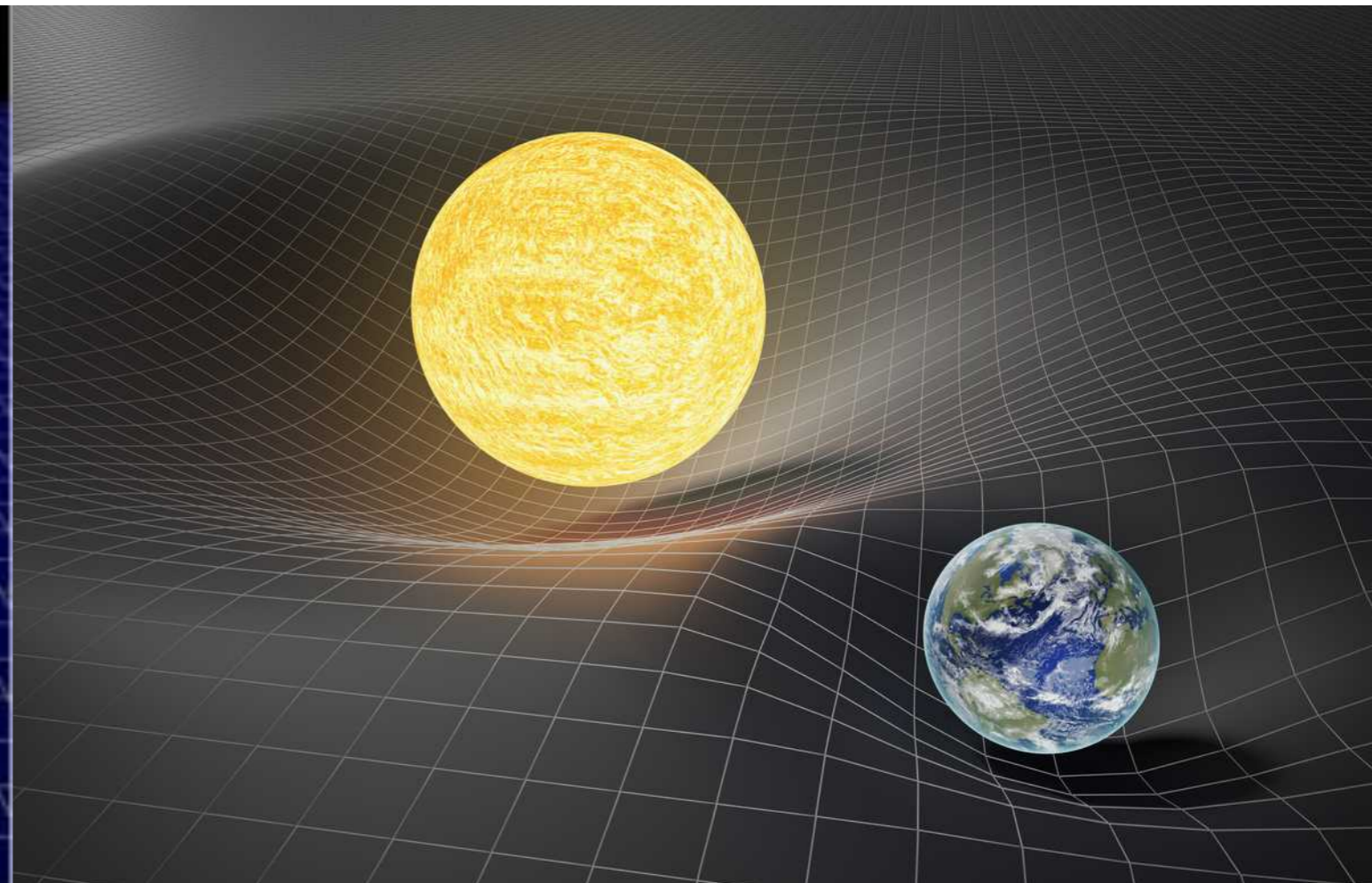
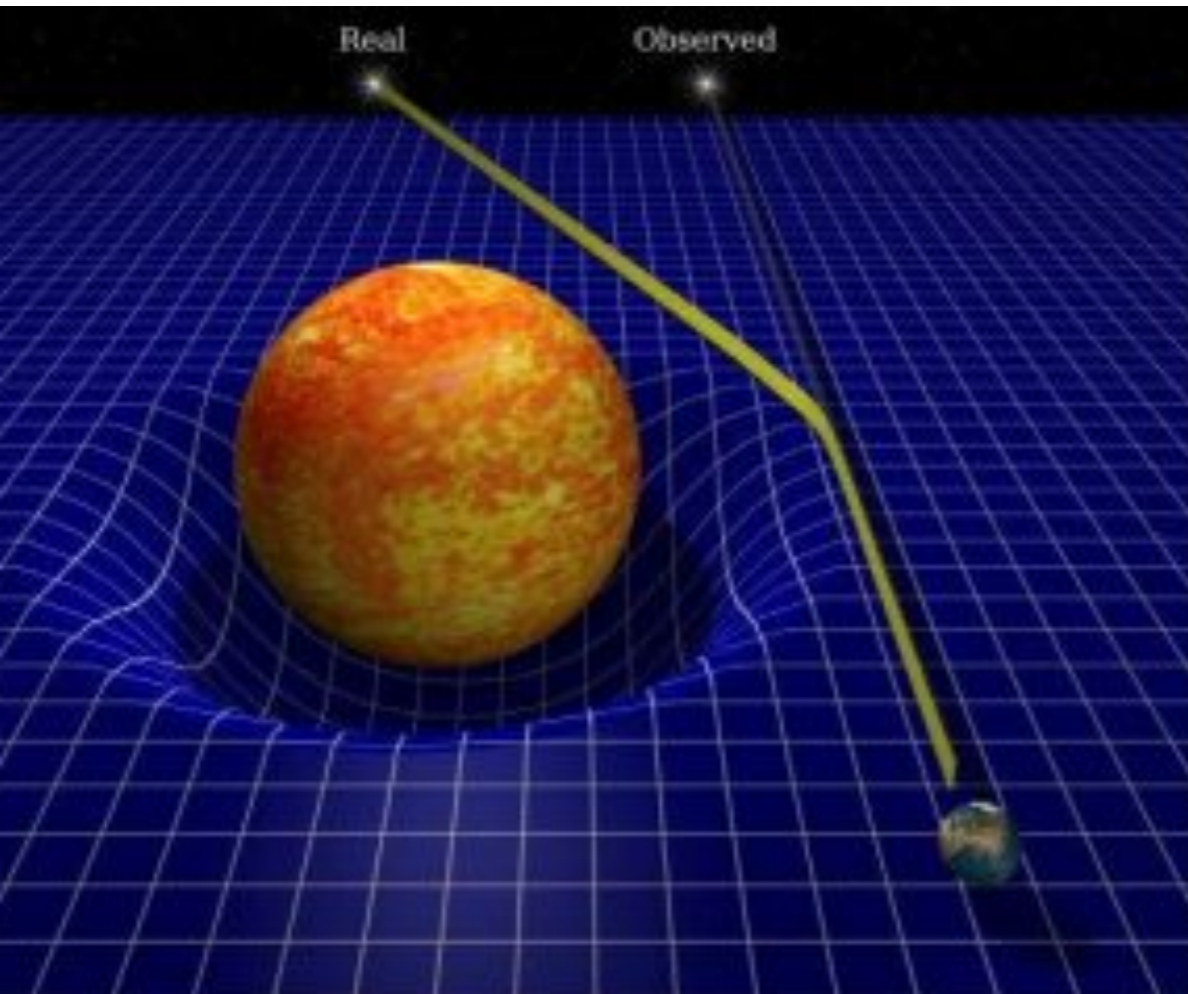
Einstein's Theory

Mass creates **curvature** of space-time

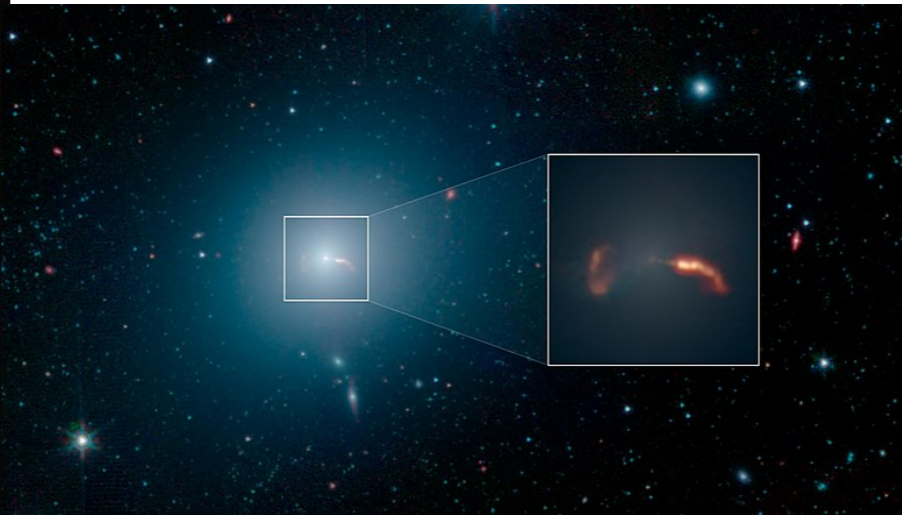
$$G = T$$

Light travels along the shortest path in the curved space

Geometry = **Energy**



Black Holes



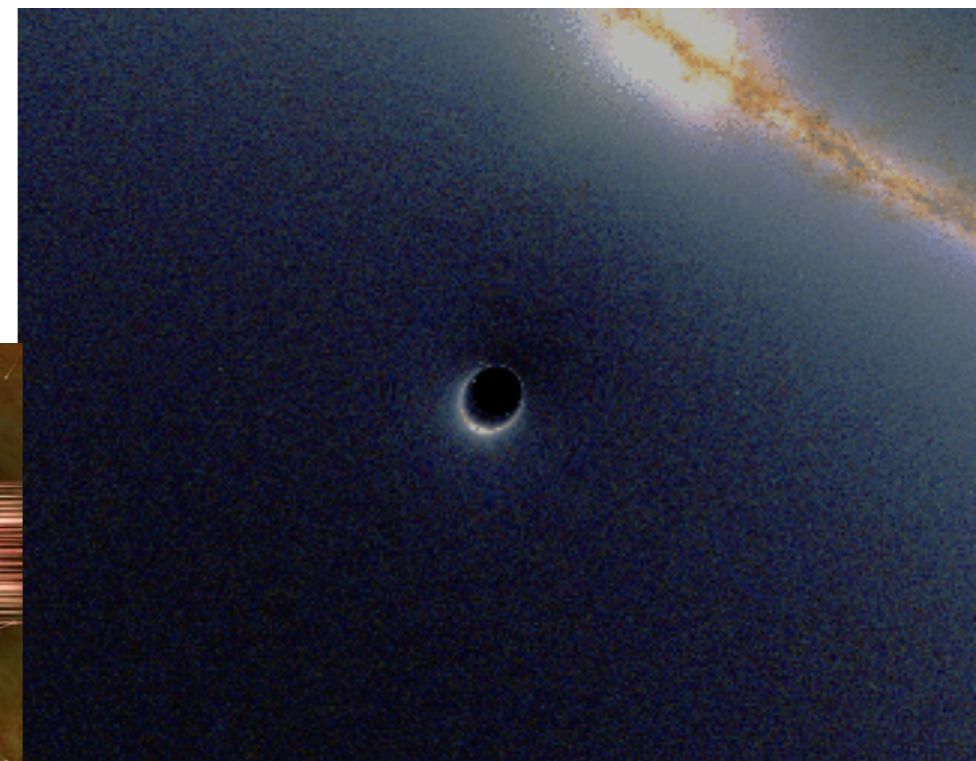
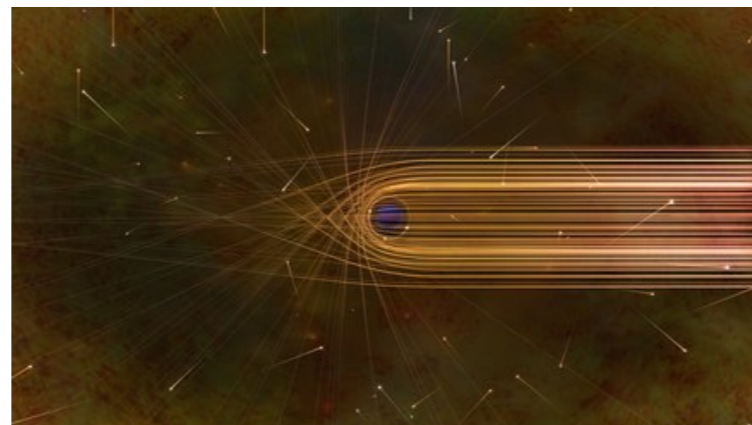
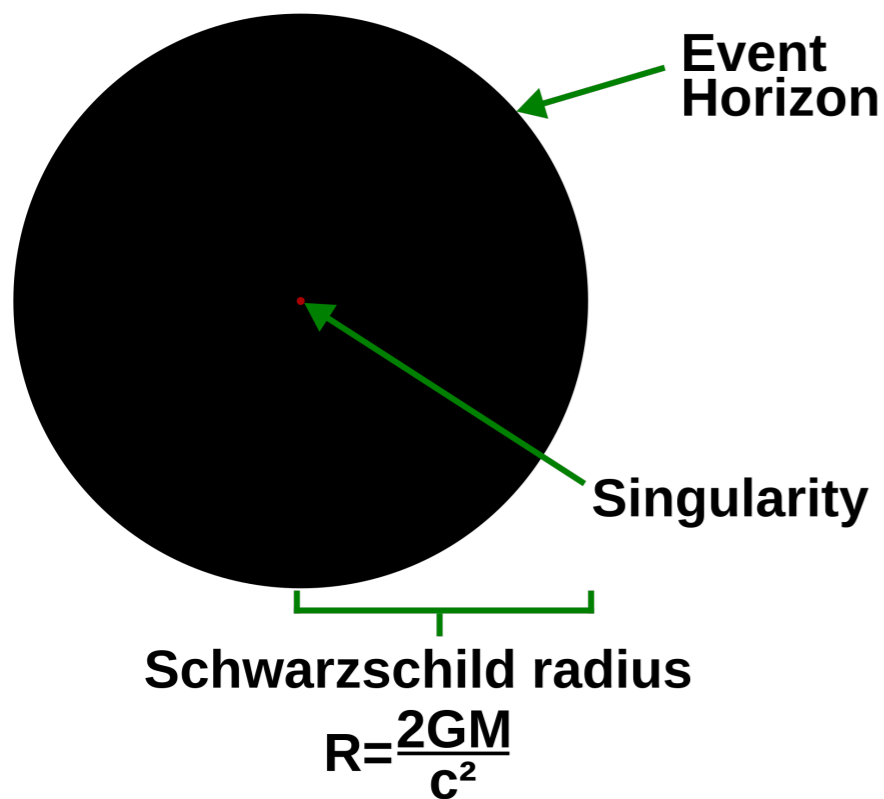
Too massive that even **light** cannot escape!

$$r_S = \frac{2G_N M}{c^2} \approx 3 \frac{M}{M_{\text{Sun}}} \text{ km}$$

Time *s l o w s d o w n* as we approach a black hole

At the horizon it **stops!**

The horizon is the **'zone of no return'**

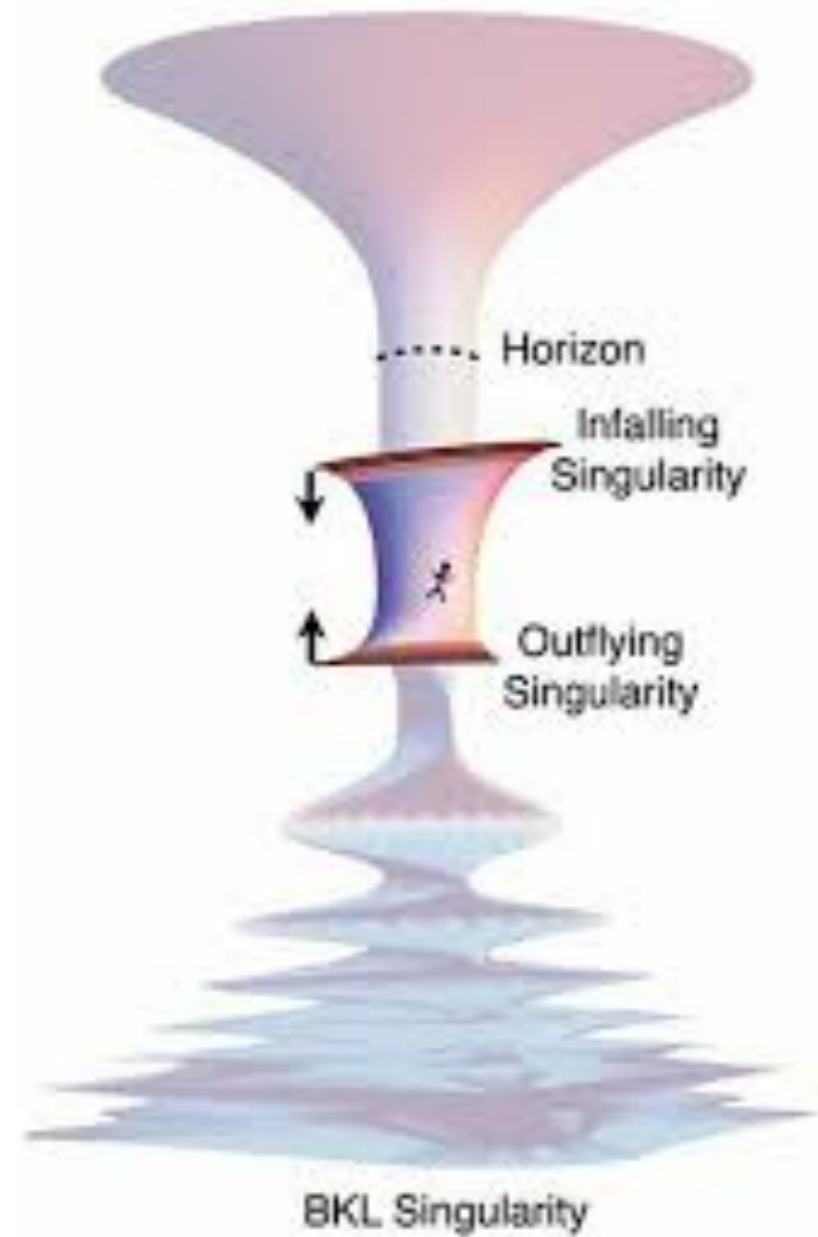
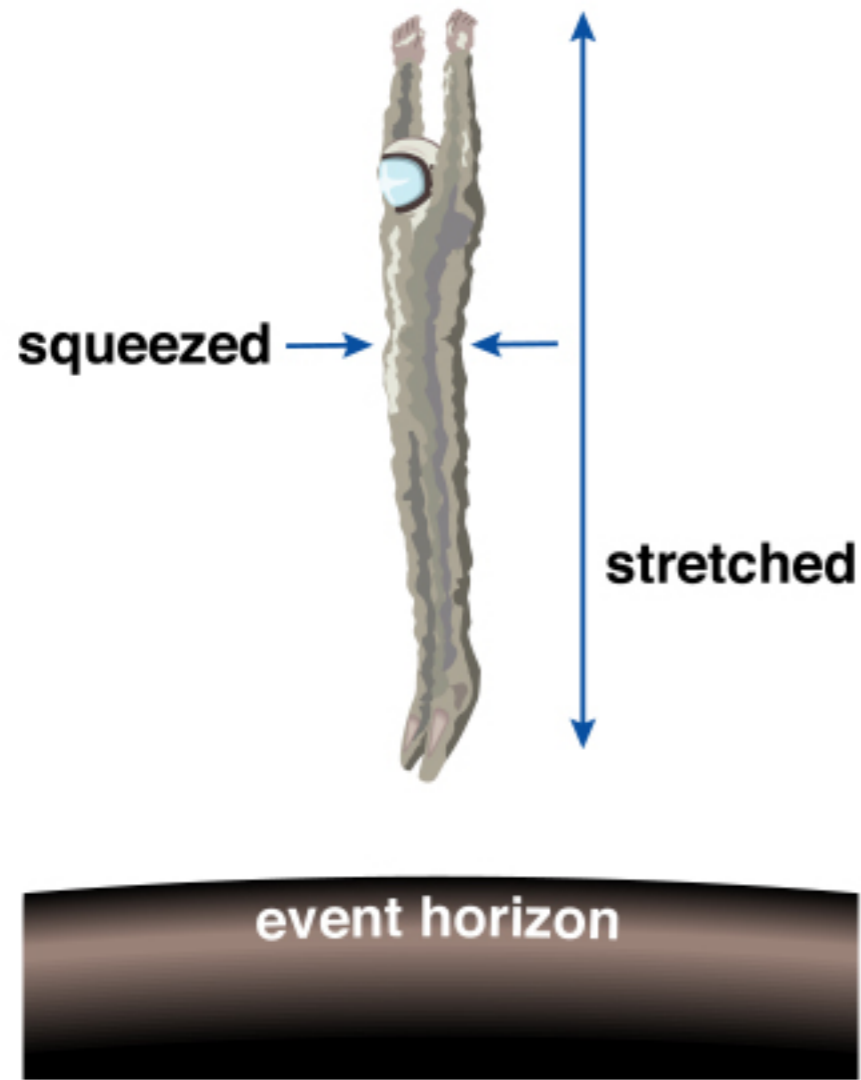


SIZE COMPARISON: THE M87 BLACK HOLE AND OUR SOLAR SYSTEM

EHT BLACK HOLE IMAGE
SOURCE: NSF



Falling into Black Hole

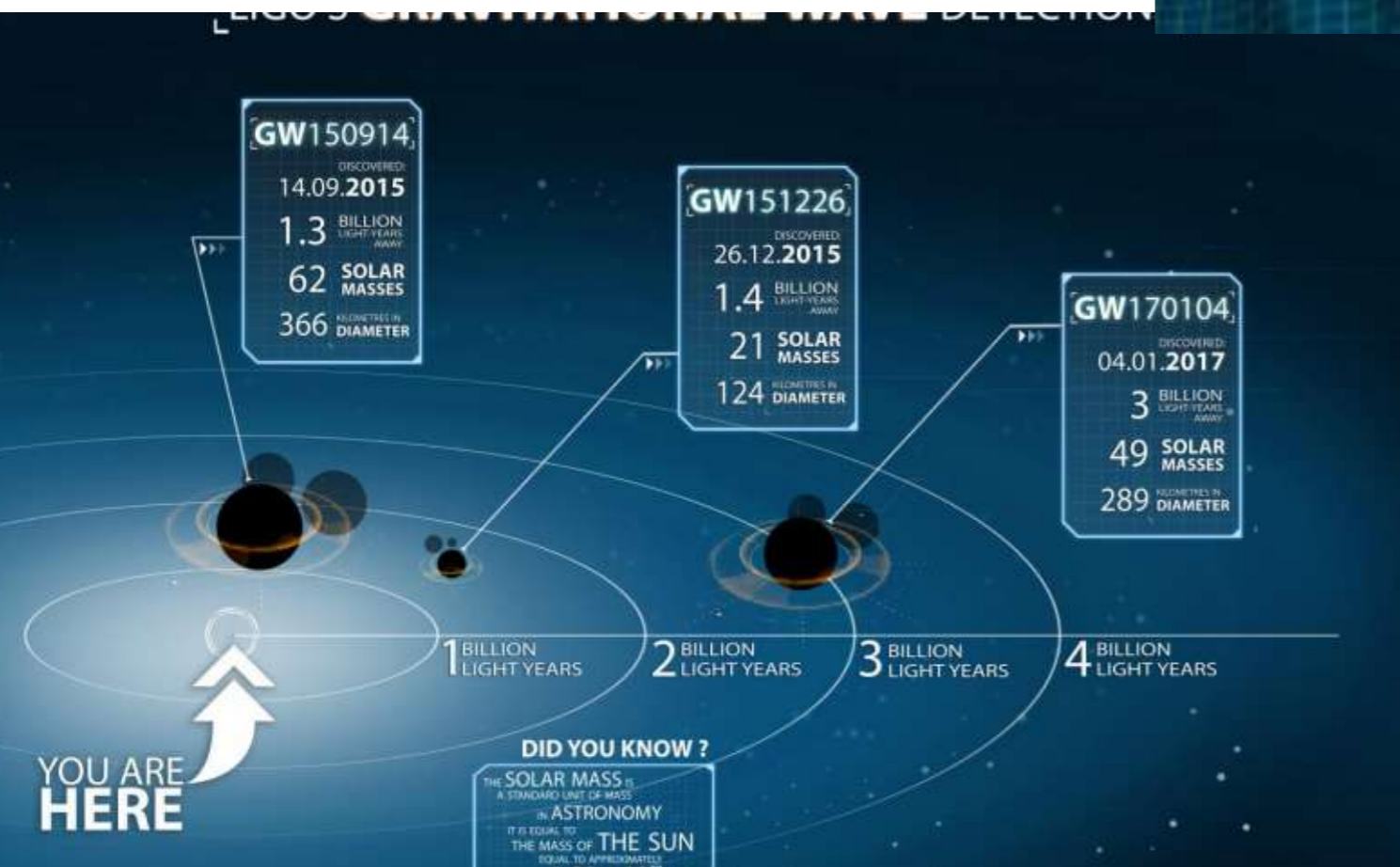
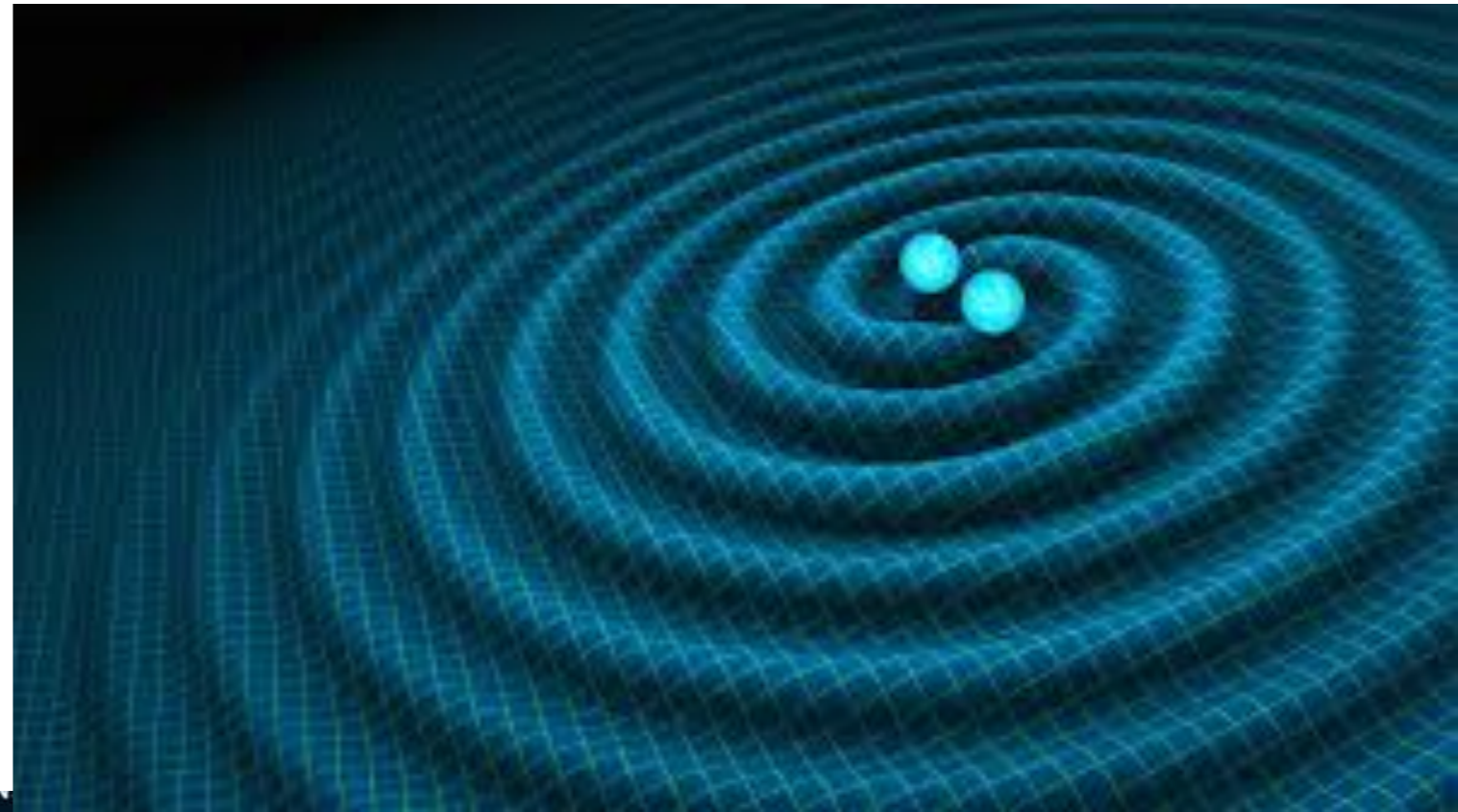


Gravitational Waves

Einstein equations

Geometry = Energy

$$G = T$$

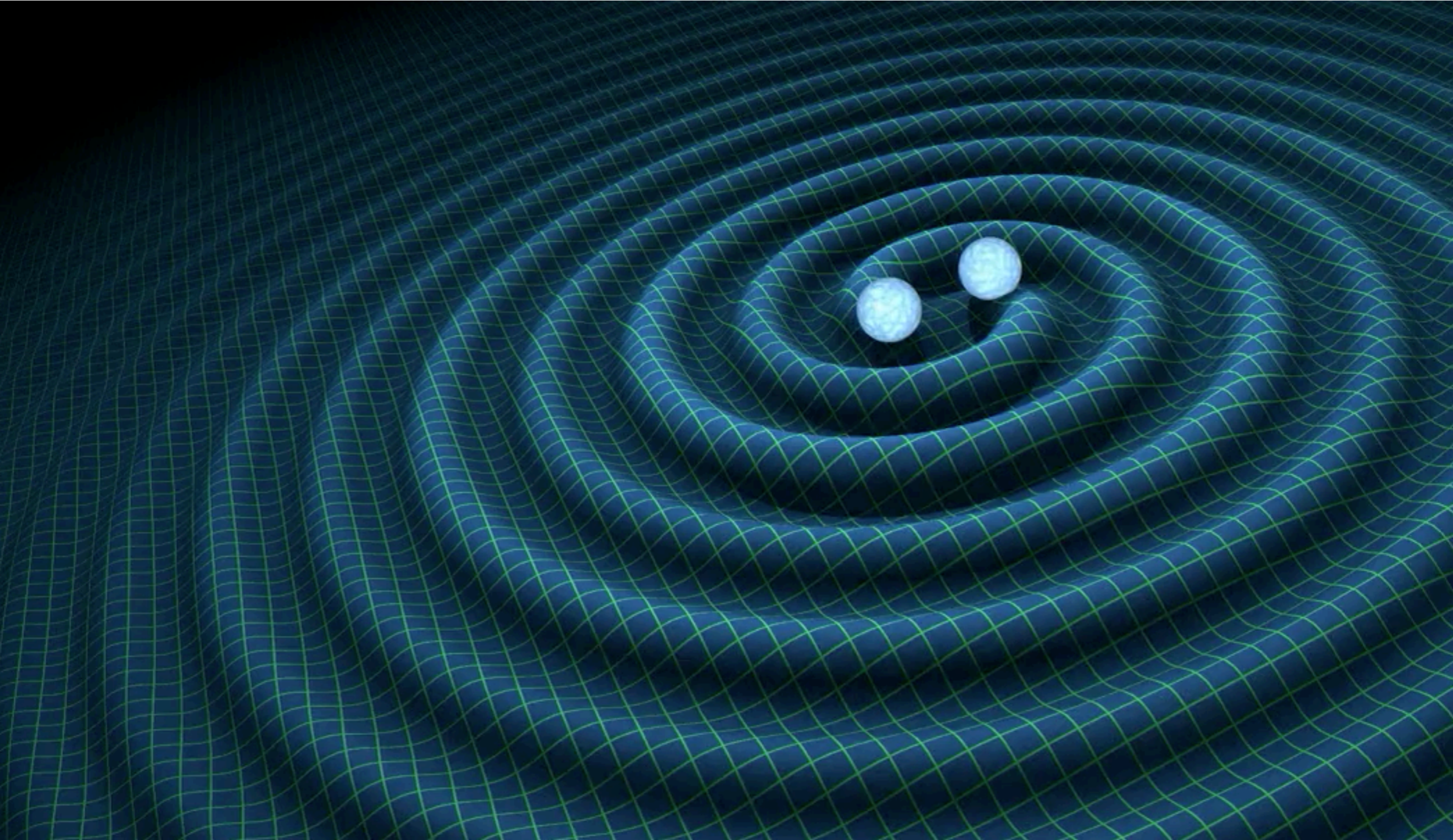


Fluctuations of the **fabric of space**

$$G = G + \delta G$$

$$T = T + \delta T$$

Waves from Binaries

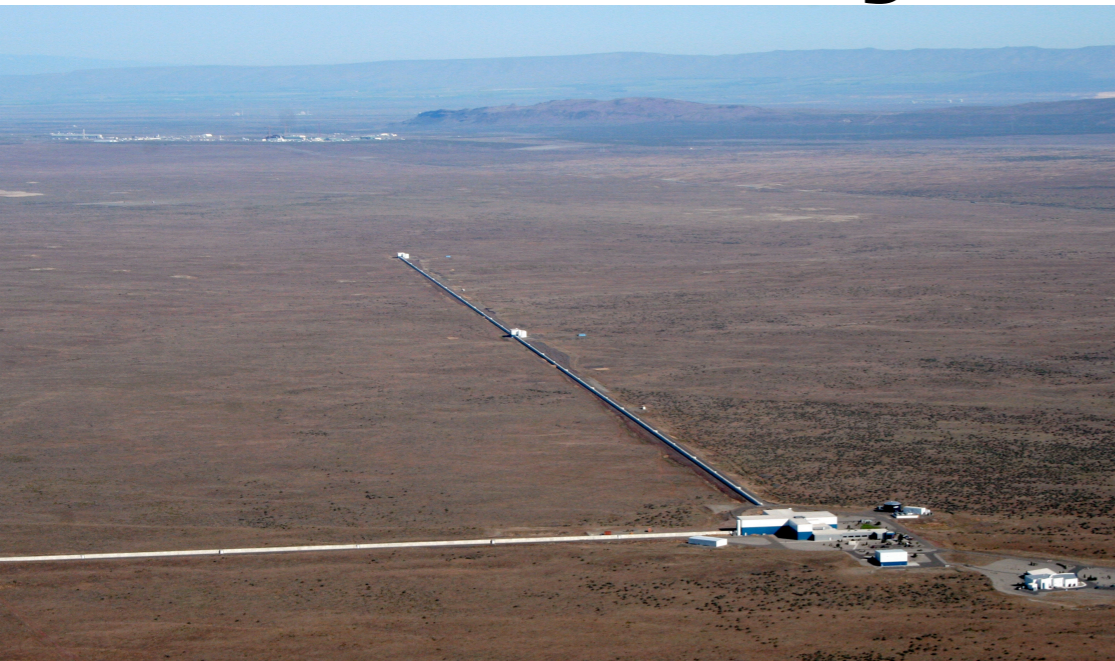


Gravity Wave



Scale of Effect Vastly Exaggerated

Gravity Wave Detectors



The Nobel Prize in Physics 2017



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Rainer Weiss
 Prize share: 1/2

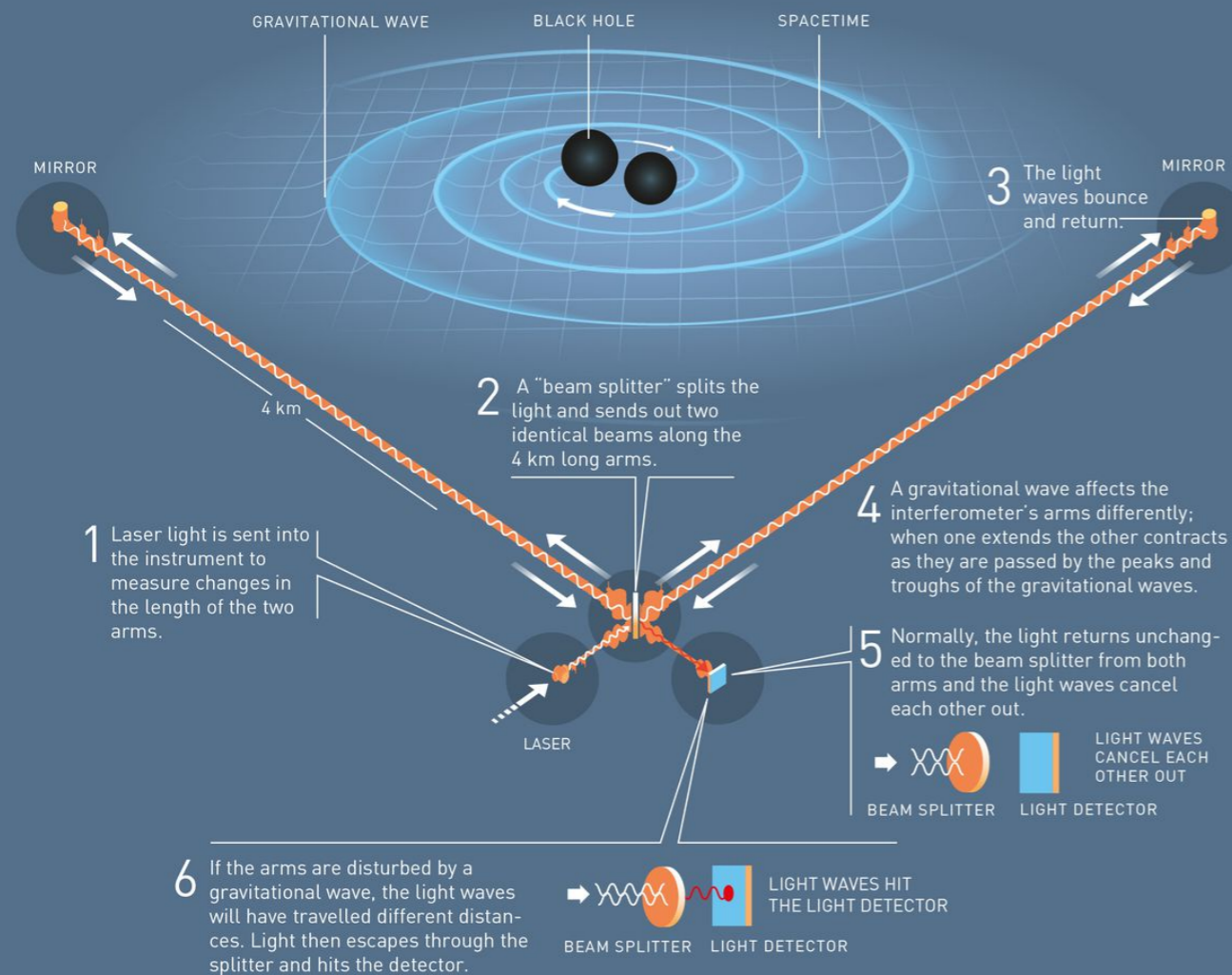


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Barry C. Barish
 Prize share: 1/4

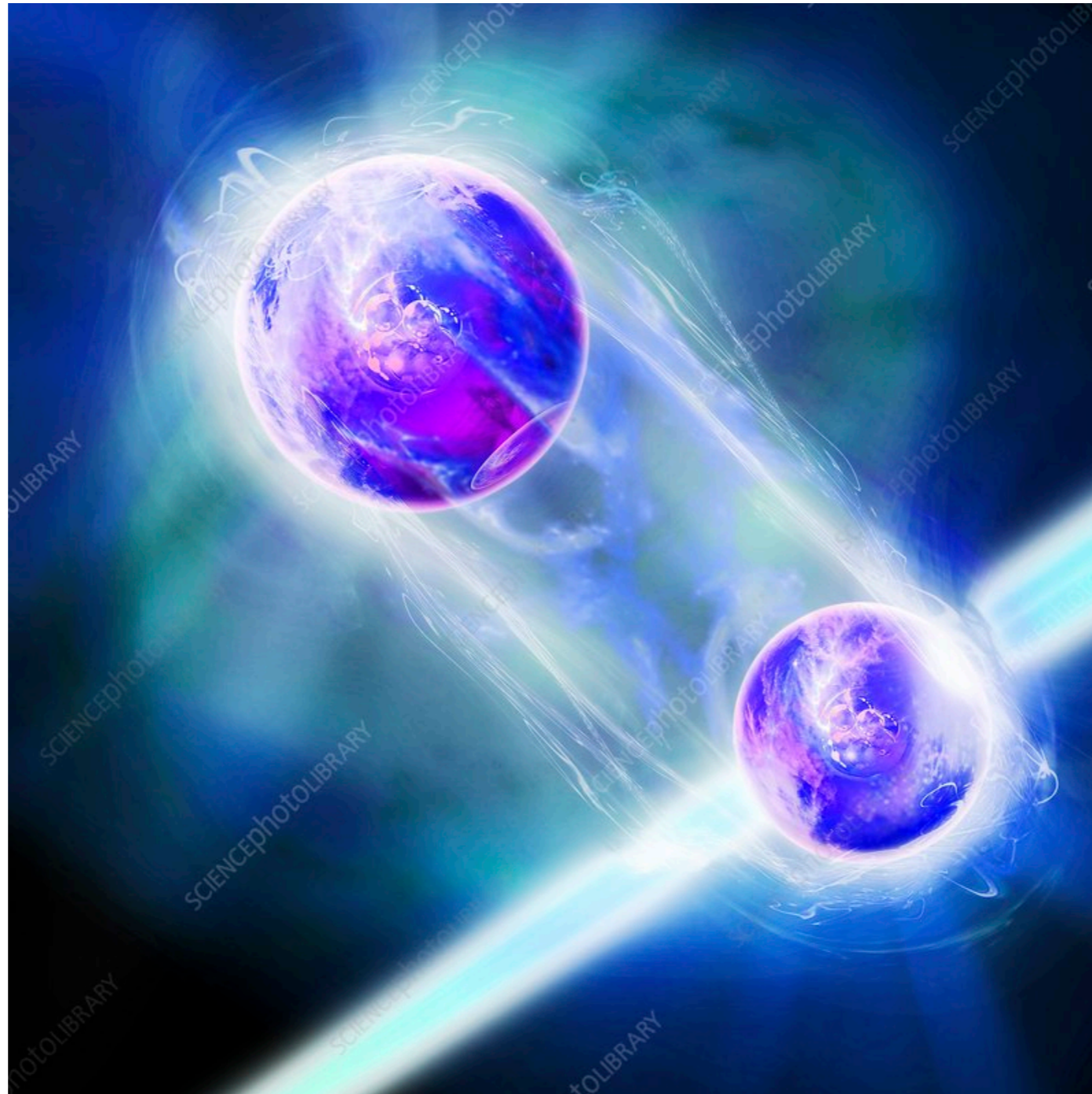


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Kip S. Thorne
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LIGO – A GIGANTIC INTERFEROMETER

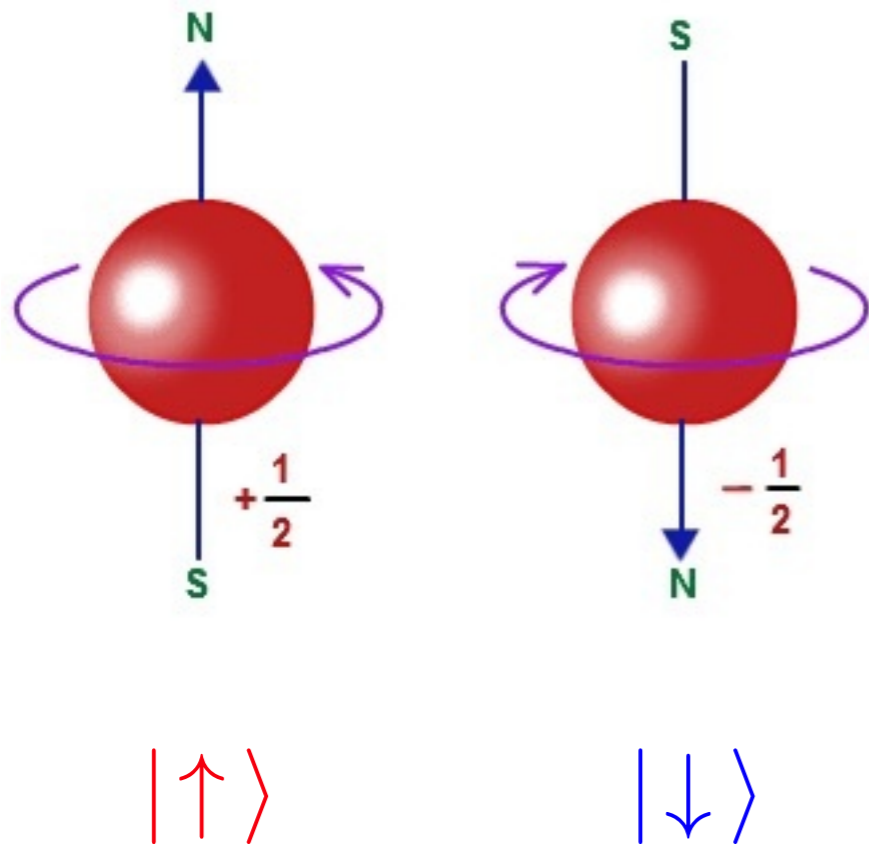


Quantum Mechanics & Entanglement



Spin

Electron can have spin up or down



According to quantum mechanics we can find an electron either with spin up or down with a probability

$$|\Psi\rangle = a|\uparrow\rangle + b|\downarrow\rangle$$

such that $a^2 + b^2 = 1$

Entanglement

Let us now take two electrons

All possible states

$$|\uparrow\rangle|\uparrow\rangle \quad |\uparrow\rangle|\downarrow\rangle \quad |\downarrow\rangle|\uparrow\rangle \quad |\downarrow\rangle|\downarrow\rangle$$

General state

$$|\Psi\rangle = c |\uparrow\rangle|\uparrow\rangle + d |\uparrow\rangle|\downarrow\rangle + e |\downarrow\rangle|\uparrow\rangle + f |\downarrow\rangle|\downarrow\rangle$$

Cannot be generally written as product

$$(a |\uparrow\rangle + b |\downarrow\rangle)(A |\uparrow\rangle + B |\downarrow\rangle)$$

If not then we say the spin 1 is **entangled** with spin 2

Examples

Not entangled

$$|\downarrow\rangle|\uparrow\rangle$$
$$|\uparrow\rangle|\uparrow\rangle + |\uparrow\rangle|\downarrow\rangle = |\uparrow\rangle(|\uparrow\rangle + |\downarrow\rangle)$$

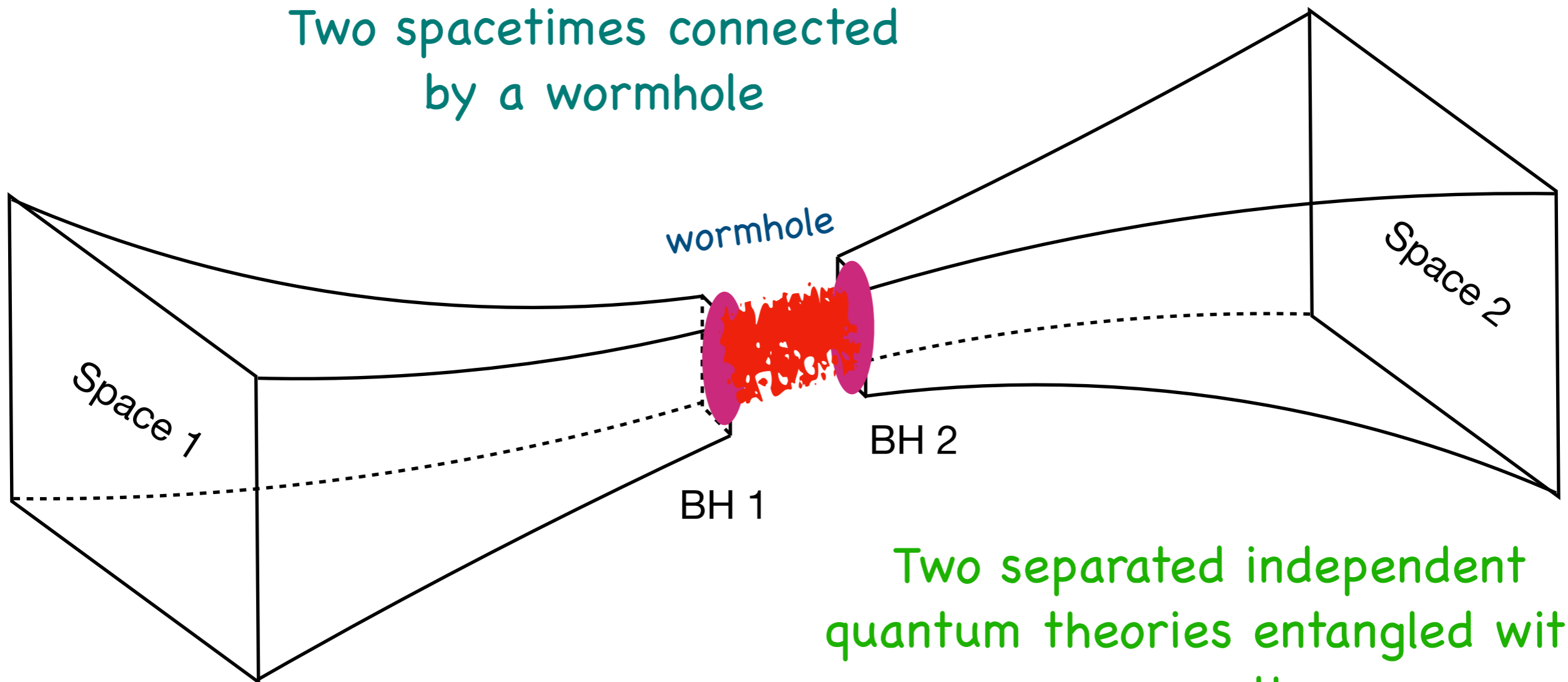
Entangled

$$|\uparrow\rangle|\downarrow\rangle + |\downarrow\rangle|\uparrow\rangle$$

$$|\uparrow\rangle|\uparrow\rangle + |\downarrow\rangle|\downarrow\rangle$$

Entanglement in Gravity

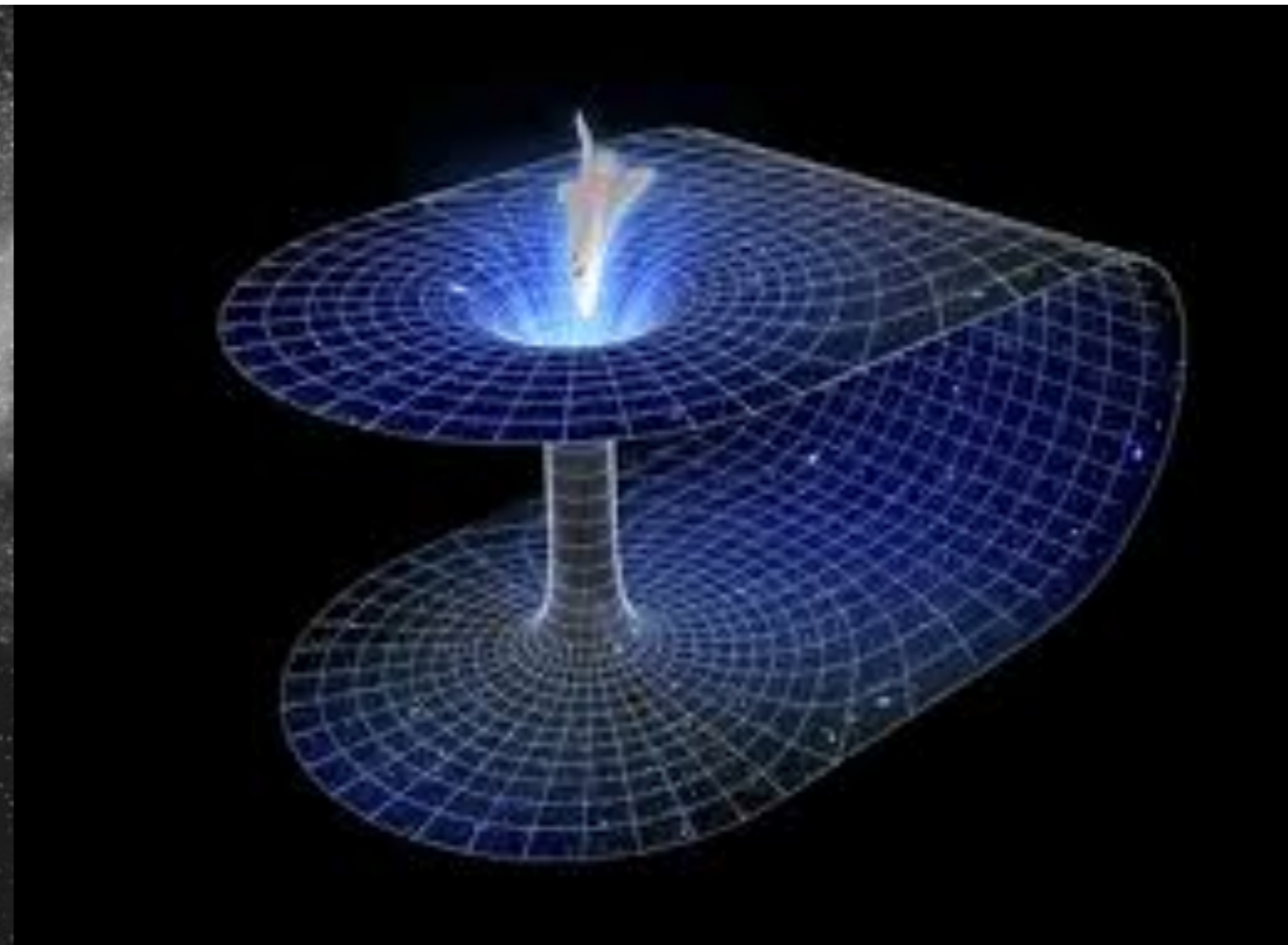
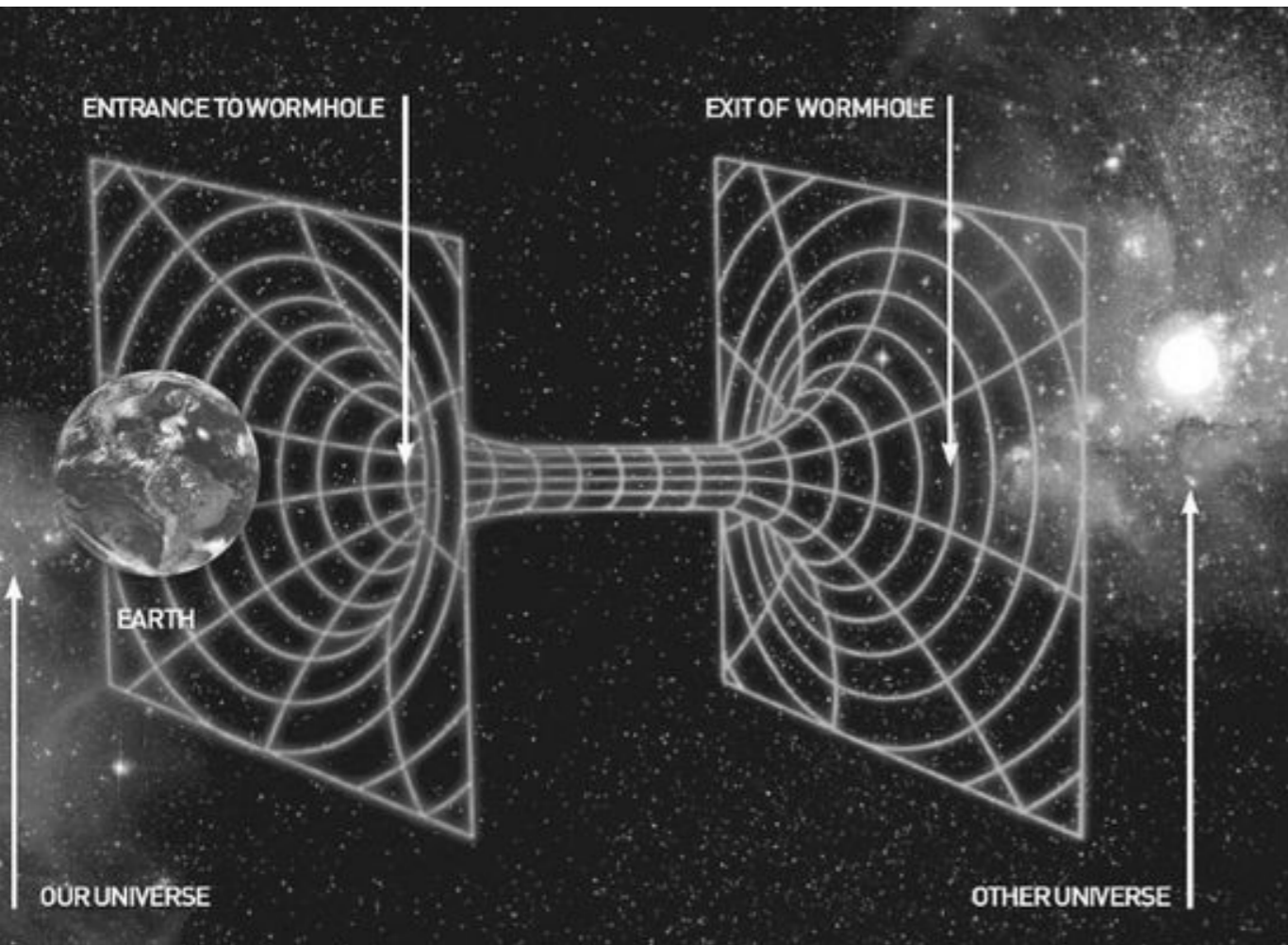
Two spacetimes connected
by a wormhole



Two separated independent
quantum theories entangled with
one another

$$|\Psi\rangle = |0\rangle_1 |0\rangle_2 + c |\uparrow\rangle_1 |\uparrow\rangle_2 + d |\uparrow\rangle_1 |\downarrow\rangle_2 + \dots$$

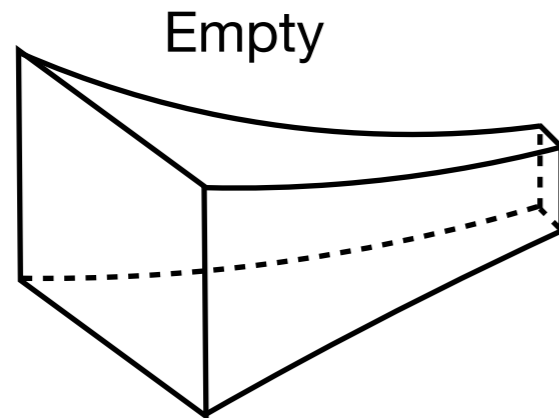
Wormholes



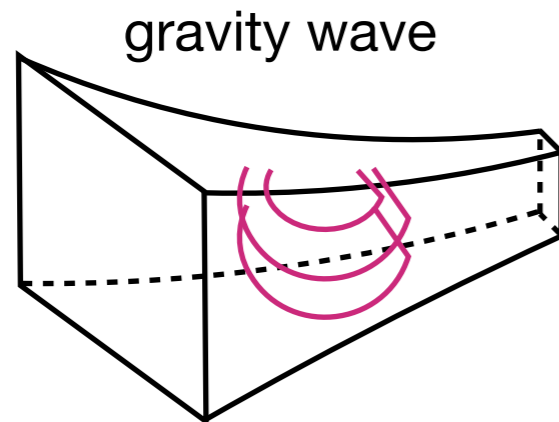
Holographic Principle

Spacetime

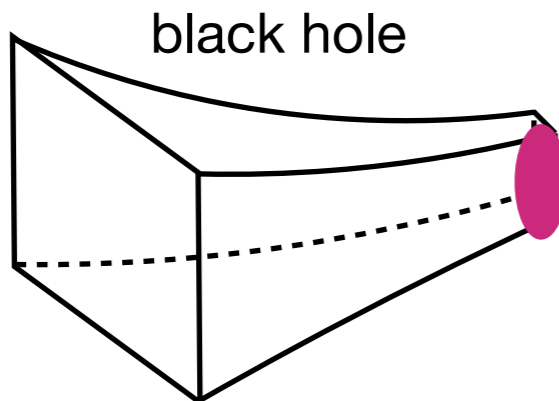
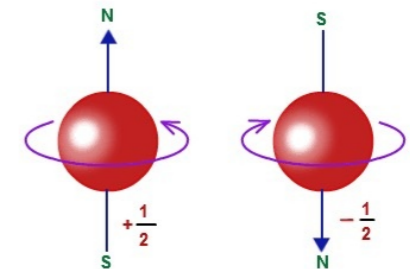
States in quantum field theory



vacuum $|0\rangle$

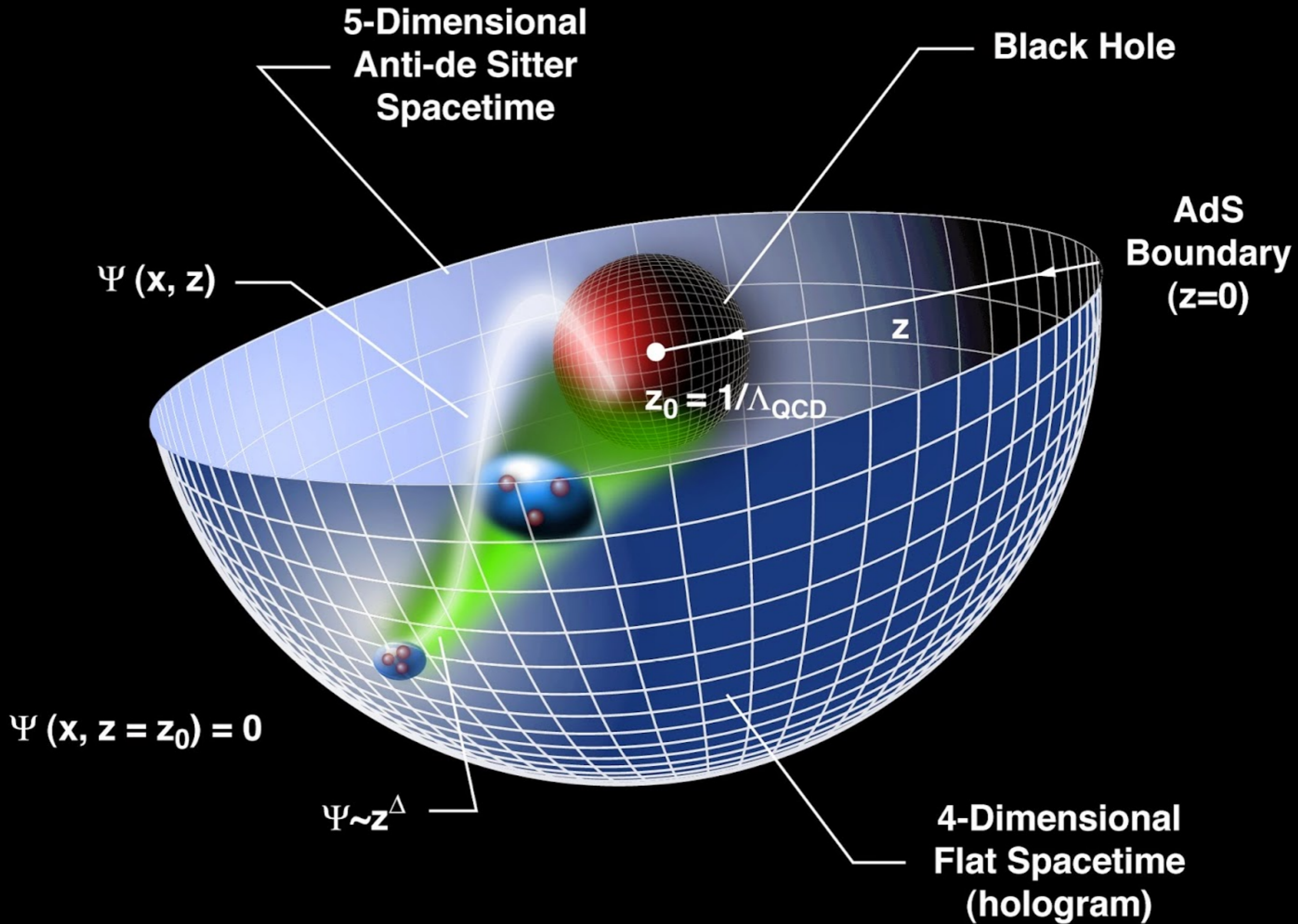


particles



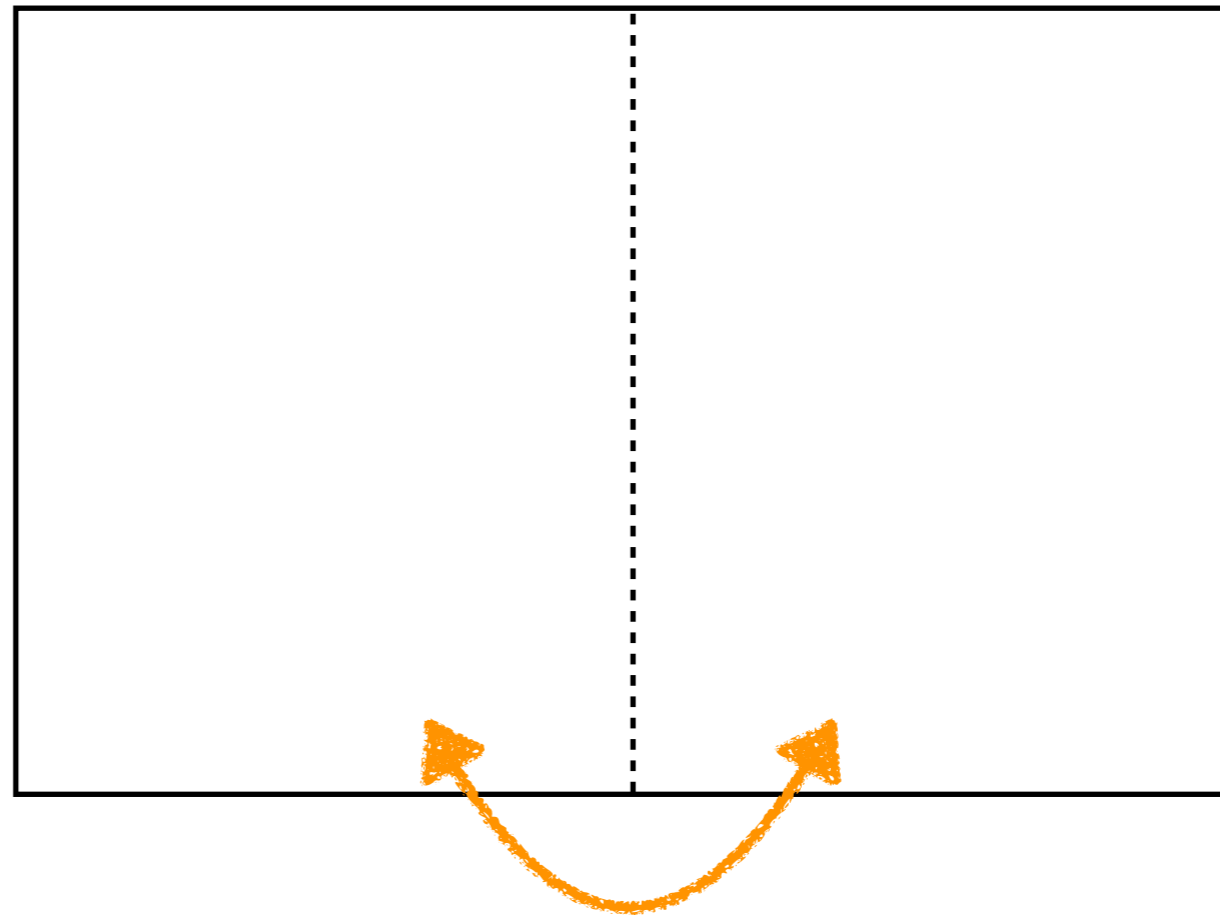
higher excitations





Entanglement in QFT

Vacuum state has a lot of entanglement already

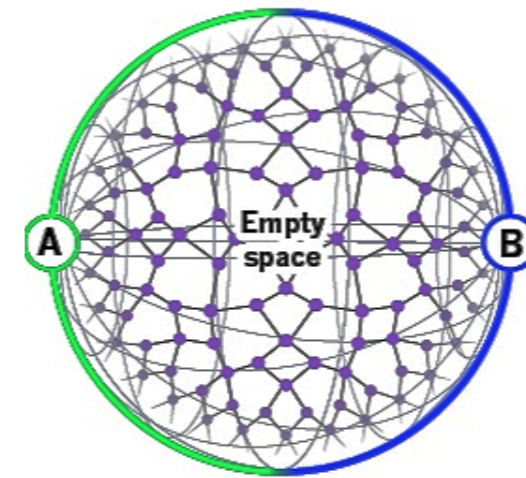
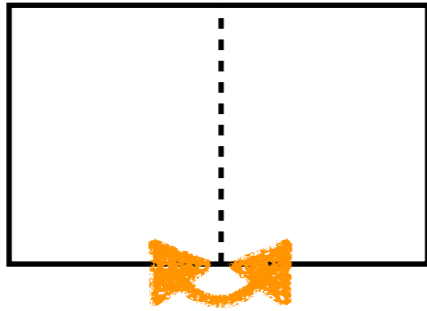


In this room electromagnetic field on one side is entangled with that on another side

Pulling space apart

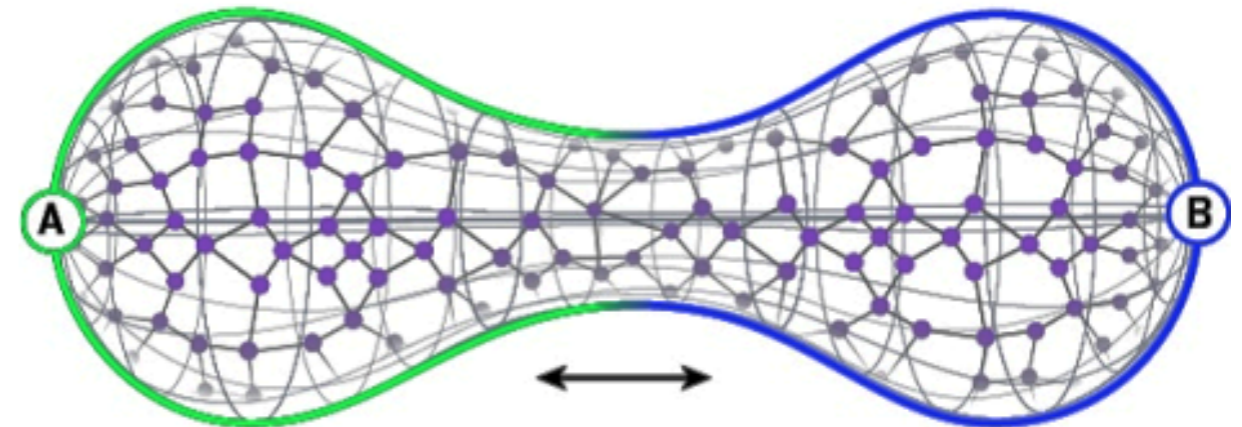
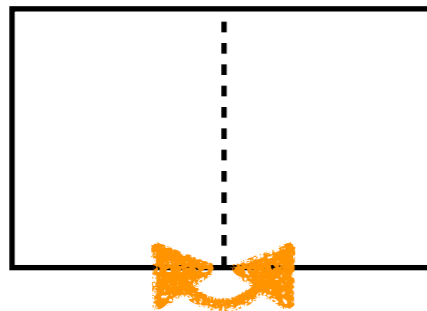
Highly entangled

$|0\rangle$



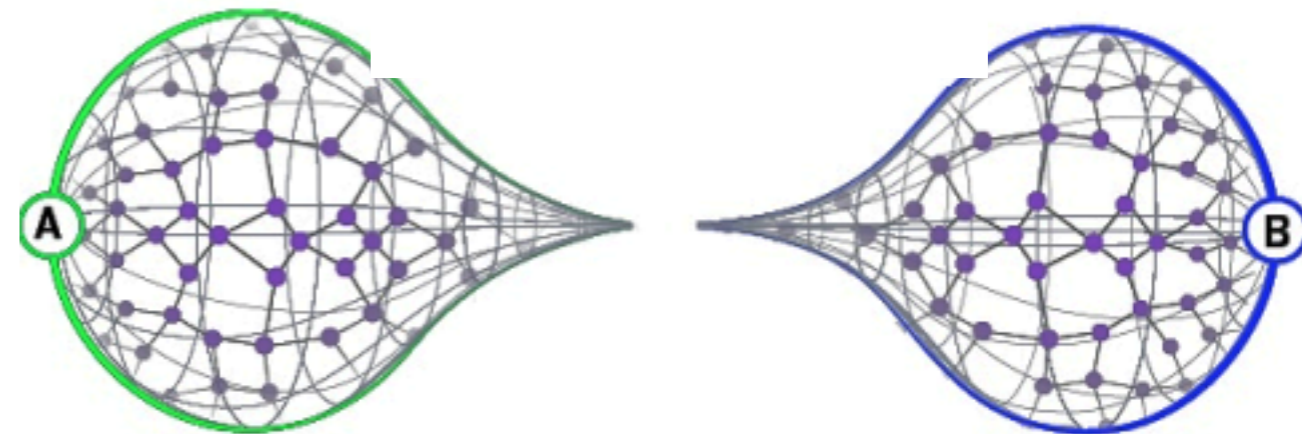
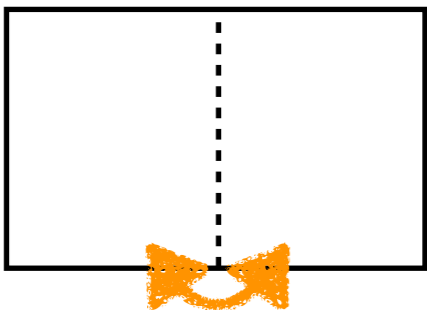
Less entangled

$|\Psi_1\rangle$



Even less entangled

$|\Psi_2\rangle$



Space via Entanglement

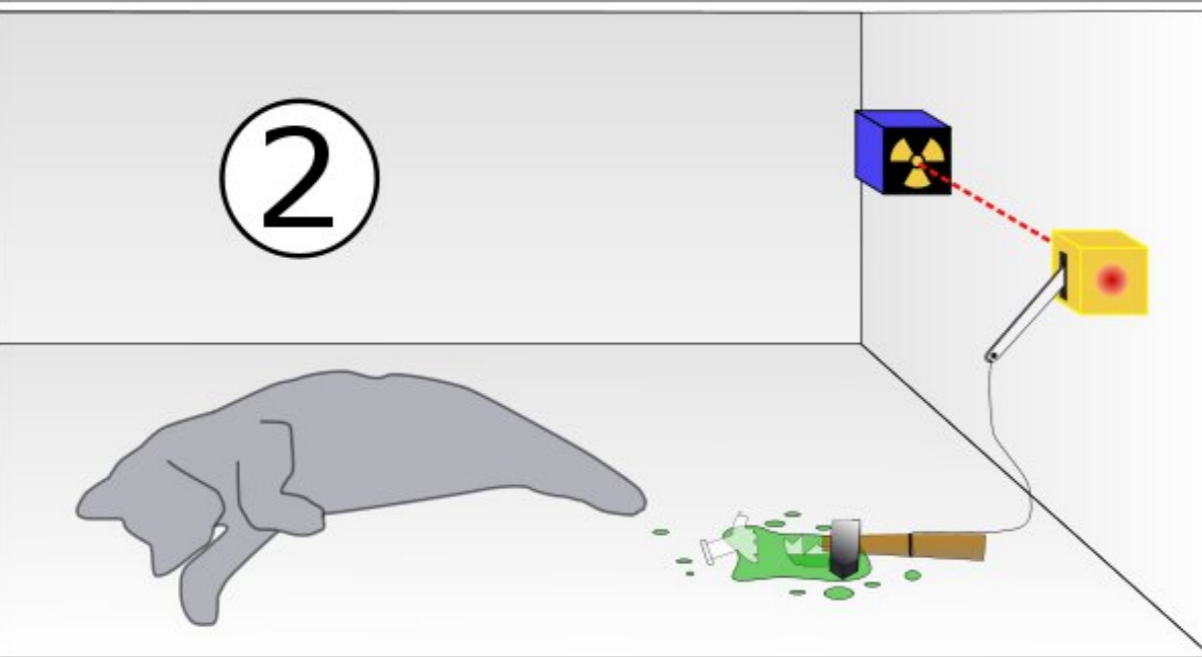
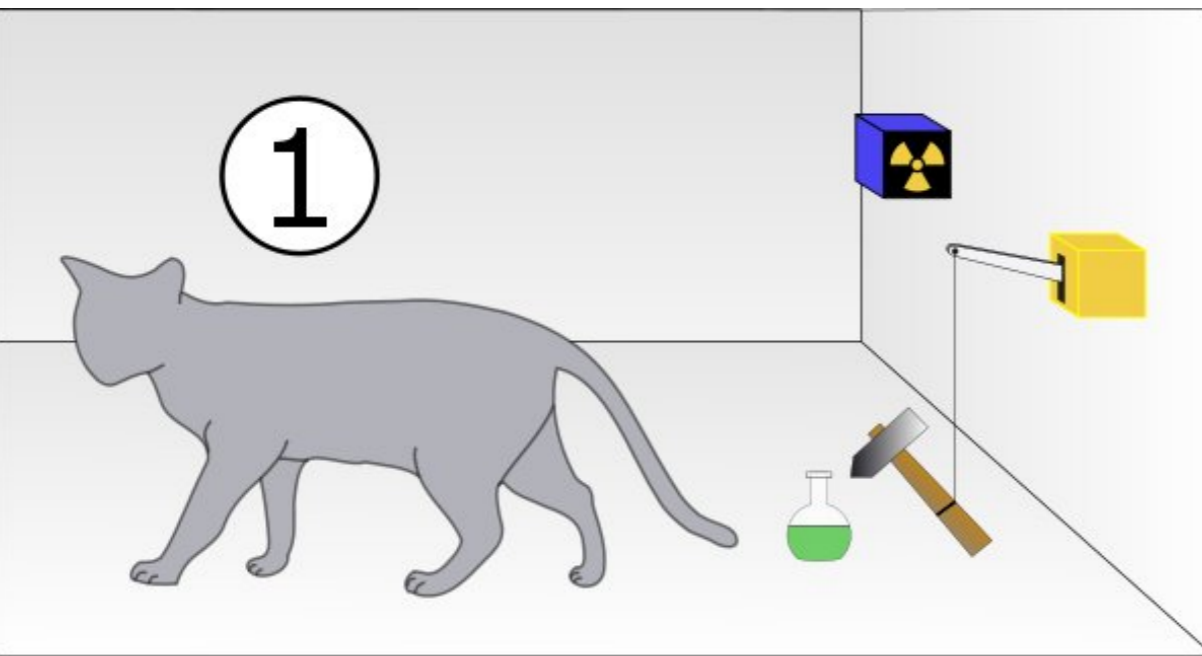
This suggests that **spacetime geometry** emerges via **entanglement** of degrees of freedom of the dual field theory!

In other words, there is no classical spacetime without entanglement



Measurement in Quantum Mechanics

Schrodinger's cat



We cannot find out till we open the box whether:

$$|\uparrow\rangle$$

Cat is alive

$$|\downarrow\rangle$$

Cat is dead

Measuring Probabilities

Not entangled

$$\begin{array}{l}
 |\downarrow\rangle|\uparrow\rangle \xrightarrow{\hspace{10em}} |\downarrow\rangle \quad p = 1 \\
 |\uparrow\rangle|\uparrow\rangle + |\uparrow\rangle|\downarrow\rangle = |\uparrow\rangle(|\uparrow\rangle + |\downarrow\rangle) \xrightarrow{\hspace{10em}} |\uparrow\rangle \quad p = 1
 \end{array}$$

Entangled

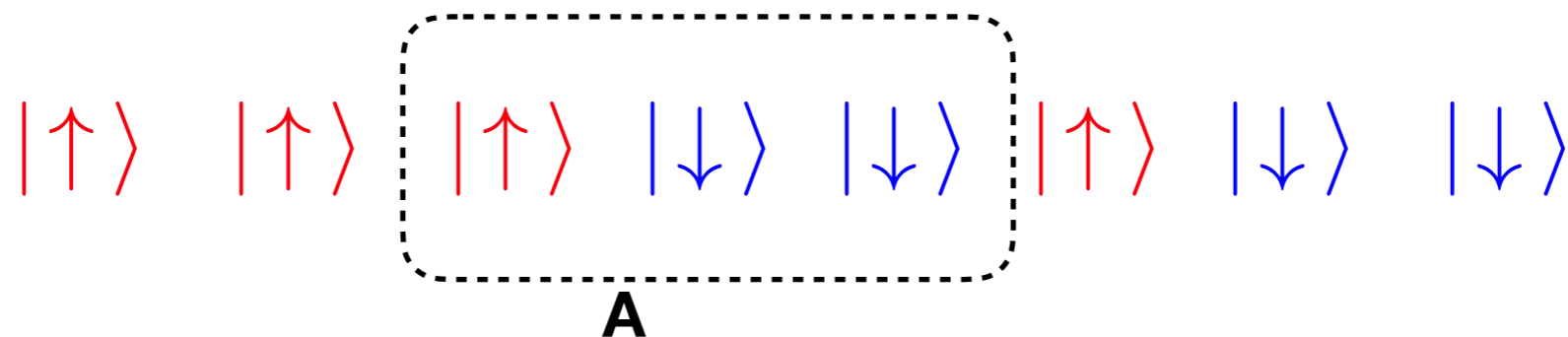
$$\begin{array}{l}
 \frac{1}{\sqrt{2}}|\uparrow\rangle|\downarrow\rangle + \frac{1}{\sqrt{2}}|\downarrow\rangle|\uparrow\rangle \xrightarrow{\hspace{10em}} |\downarrow\rangle \quad p = 1/2 \\
 \frac{1}{\sqrt{2}}|\uparrow\rangle|\uparrow\rangle + \frac{1}{\sqrt{2}}|\downarrow\rangle|\downarrow\rangle \xrightarrow{\hspace{10em}} |\uparrow\rangle \quad p = 1/2
 \end{array}$$

Less entangled

$$\begin{array}{l}
 \sqrt{\frac{9}{10}}|\uparrow\rangle|\downarrow\rangle + \sqrt{\frac{1}{10}}|\downarrow\rangle|\uparrow\rangle \xrightarrow{\hspace{10em}} |\downarrow\rangle \quad p = 1/10 \\
 \sqrt{\frac{9}{10}}|\uparrow\rangle|\uparrow\rangle + \sqrt{\frac{1}{10}}|\downarrow\rangle|\downarrow\rangle \xrightarrow{\hspace{10em}} |\uparrow\rangle \quad p = 9/10
 \end{array}$$

Entanglement Entropy

Entropy of the subsystem A counts the total number of states



$|\Psi_i\rangle$ has probability p_i

Entanglement of the subsystem A with the rest of the system can be quantified by the entropy of the ensemble when we count only states of the subsystem

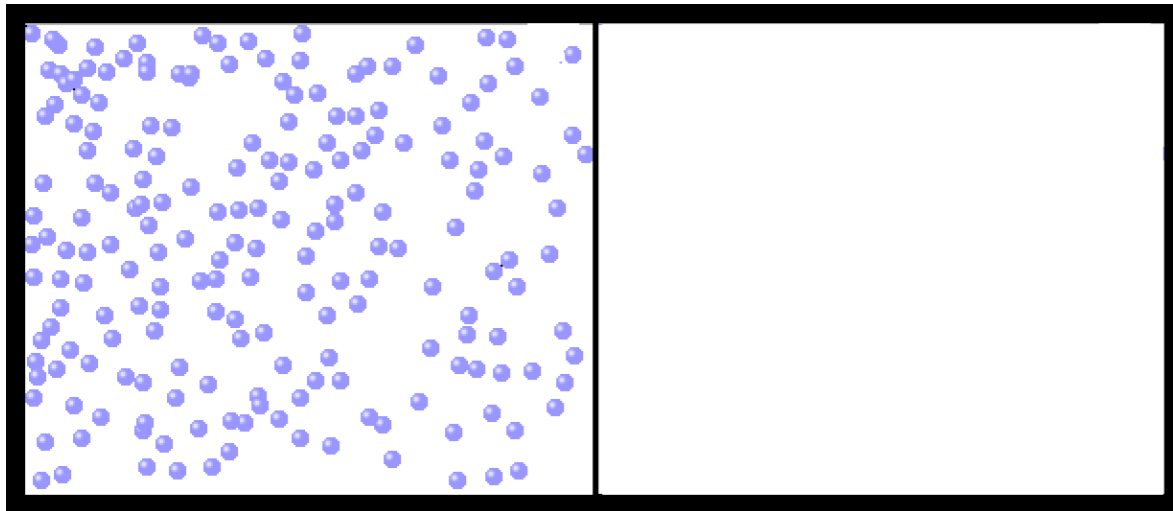
Entropy and Temperature

Entropy counts the (logarithm of) the number of states a system can have

$$S = k \log \Omega$$

Discrete in quantum statistics, continuous in classical

The second law states that entropy never decreases in a thermodynamical system



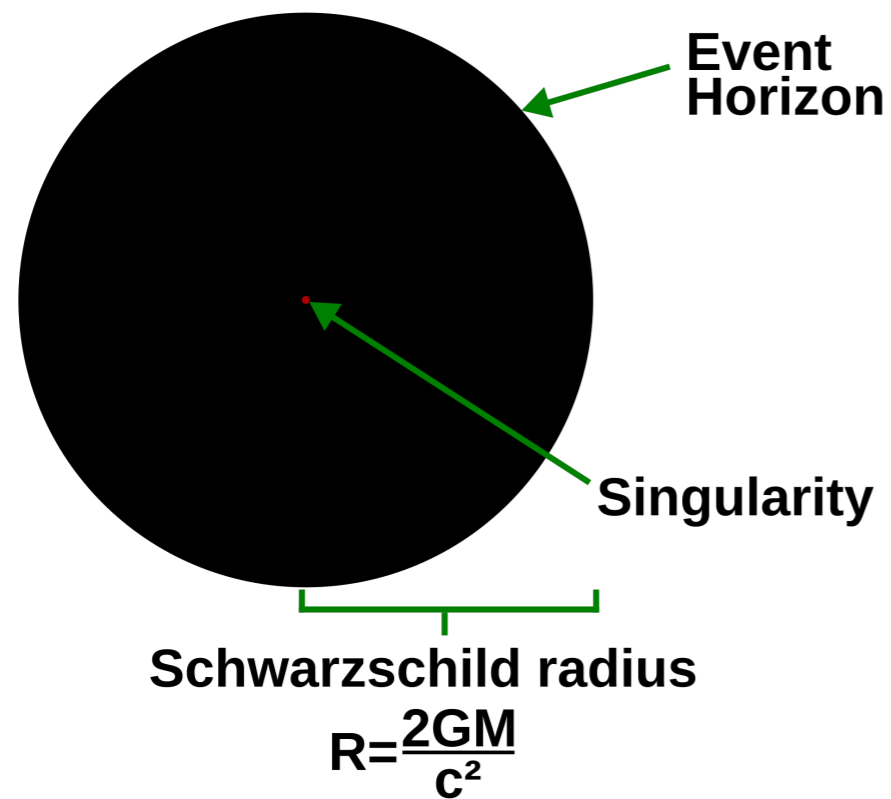
Temperature $\frac{1}{T} = \frac{\Delta S}{\Delta E}$

The Second Law of Thermodynamics, Stated as the Law of Entropy

The total entropy (or microscopic disorganization) of all the participants in any physical process cannot decrease during that process, but it can increase.

Black Hole Entropy

Entropy $S = \frac{\text{Area of the BH horizon}}{4}$



BH formula provides the (entanglement) entropy for a holographically dual theory at Hawking temperature

$$\text{Energy} = \frac{\text{Entropy}}{\text{Temperature}}$$

$$E = mc^2$$

Black Hole Radiation

Stephen Hawking realized that black holes should be considered as thermodynamical objects so they should have **temperature**

Entropy $S = \frac{\text{Area of the BH horizon}}{4}$

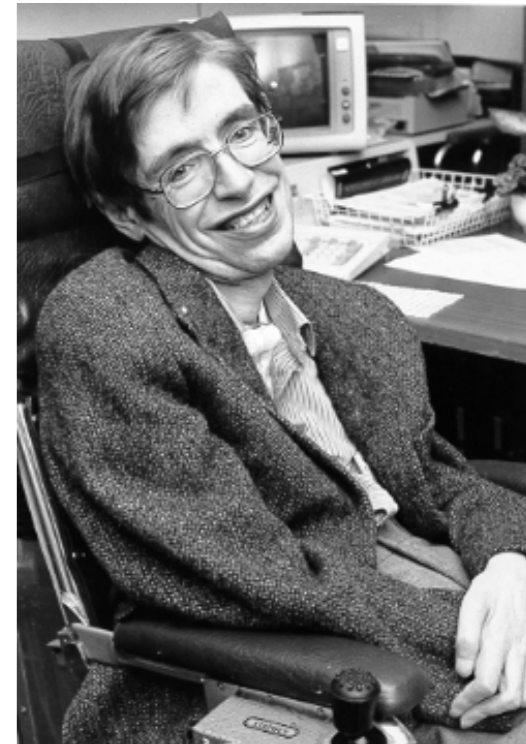
When an object falls into the BH it causes the horizon to expand slightly, which causes the changes in the entropy and hence defines the temperature

$$T_H = \frac{1}{8\pi M}$$

A body with nonzero temperature radiates (photons) so it loses energy into the outer space

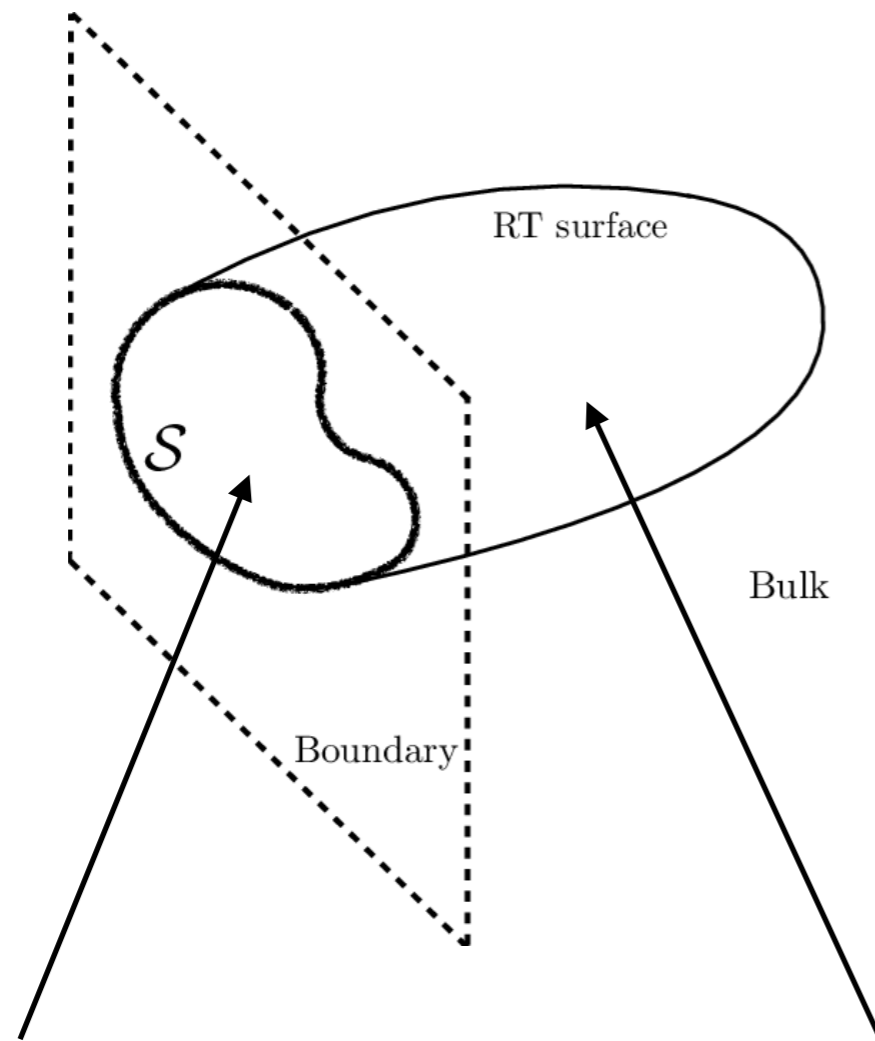
One can estimate how quickly a BH will have evaporated completely. Those times are gigantic: a solar mass black hole will take about 10^{64} years

However, tiny BHs which could exist in principle may be evaporating in the Universe right now as we speak !

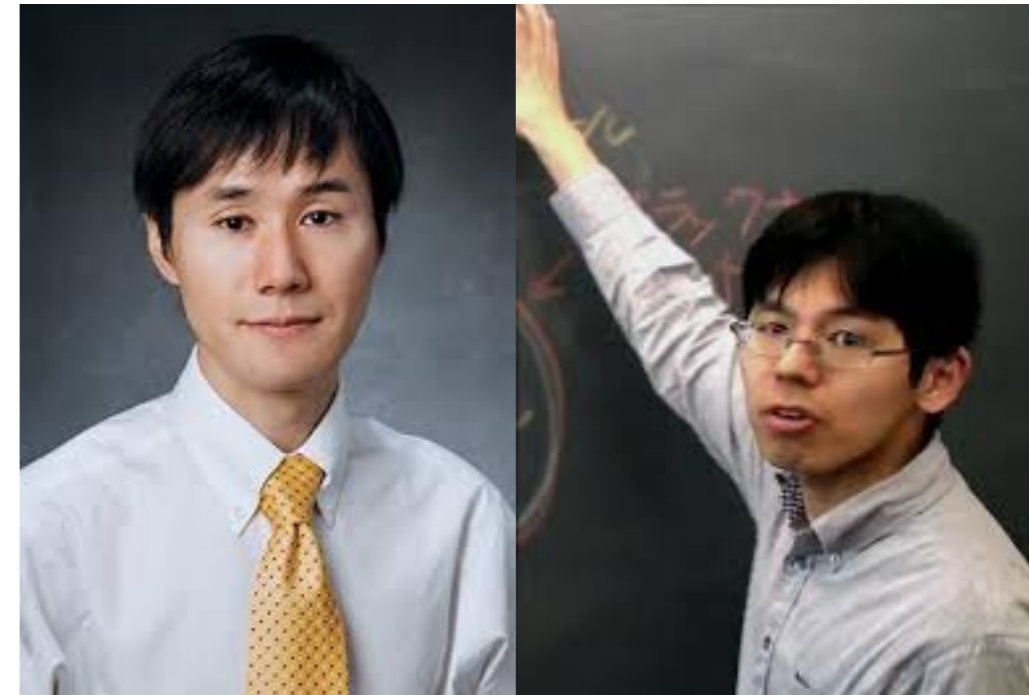


Entanglement in Holography

Entropy of any QFT subsystem (i.e. entanglement entropy) has a geometric interpretation in the dual gravity theory



Entropy of this region = Area of the minimal surface



Ryu

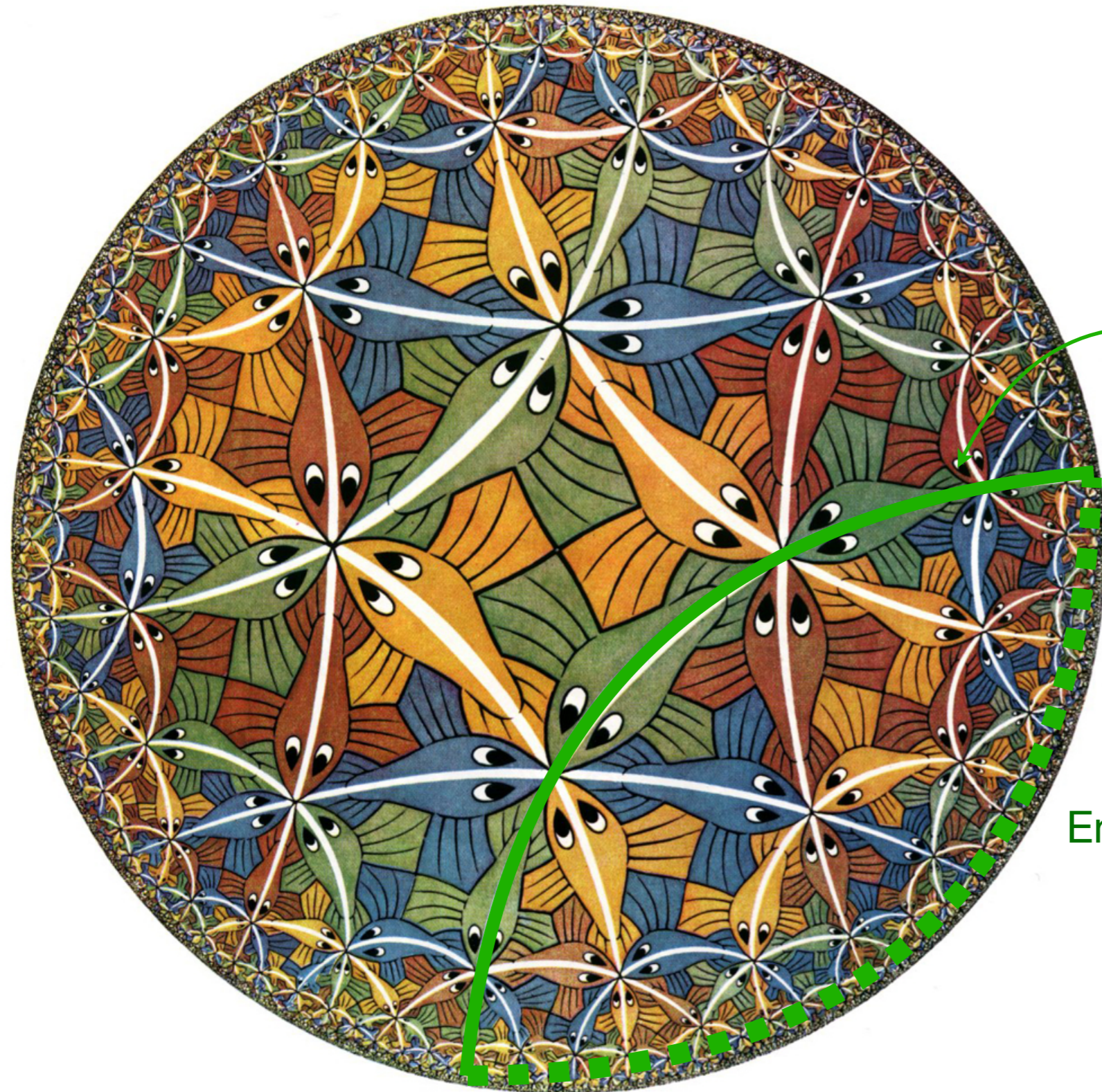
Takayanagi

Entropy vs Area

Anti-deSitter space

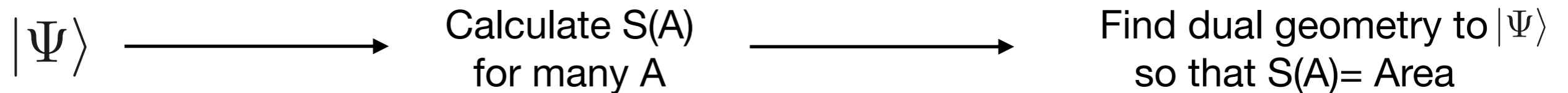
Shortest path through the bulk which has the same length as the boundary

Entanglement entropy of the boundary



From Entanglement to Geometry

One can now calculate entanglement entropy for any region of the boundary theory. We then *declare* that this **entropy** should be equal to the **area** of some region in the bulk of the dual spacetime. Thus by studying **entanglement** we can try to deduce **geometry**.



$S(A) + S(B) \geq S(AB)$ \longrightarrow These geometries should satisfy Einstein equations

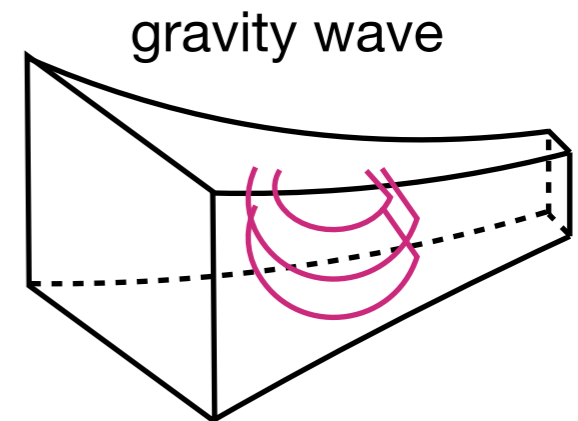
$$\Delta S(A) = \frac{\Delta \langle E \rangle}{T}$$

Perturbing States

Set of entropy constraints for all regions A on the boundary

Some math ...

Geometry must satisfy linearized Einstein equations!
(gravitational waves!)



感谢您的关注

我们在成都玩得很开心!