

UCB Math 228A, Fall 2014: Homework Set 4

Due Nov. 3, 2014

Code Submission: E-mail all requested and supporting MATLAB files to Luming at lwang@berkeley.edu as a zip-file named lastname.firstname_4.zip, for example luming-wang_4.zip.

1. The *Cash-Karp method* below is a 4/5-order embedded Runge-Kutta scheme. Recall that the vectors b^T and \hat{b}^T in the last two rows are used to produce two solution updates,

$$u_{n+1} = u_n + h \sum_{j=1}^s b_j k_j, \quad \hat{u}_{n+1} = u_n + h \sum_{j=1}^s \hat{b}_j k_j, \quad (1)$$

where \hat{u} is used only for estimation of the local truncation error $\tau \approx \|\hat{u}_{n+1} - u_{n+1}\|_\infty$.

| | | | | | | |
|------|------------|---------|-------------|--------------|-----------|----------|
| 0 | | | | | | |
| 1/5 | 1/5 | | | | | |
| 3/10 | 3/40 | 9/40 | | | | |
| 3/5 | 3/10 | -9/10 | 6/5 | | | |
| 1 | -11/54 | 5/2 | -70/27 | 35/27 | | |
| 7/8 | 1631/55296 | 175/512 | 575/13824 | 44275/110592 | 253/4096 | |
| | 2825/27648 | 0 | 18575/48384 | 13525/55296 | 277/14336 | 1/4 |
| | 37/378 | 0 | 250/621 | 125/594 | 0 | 512/1771 |

Implement a MATLAB function `rkck.m` with the syntax

```
function [t,u]=rkck(f,tlim,u0,abstol,hmin,hmax)
```

for solving ODEs using the Cash-Karp scheme with step size control. The input parameters are the right-hand side function $f(t,u)$, the start and end times $tlim=[t_0,t_1]$, the initial solution u_0 , the absolute tolerance $abstol=\delta$, and the smallest and largest acceptable timesteps $hmin$, $hmax$. The outputs are the times t and the solutions u at the actual timesteps (including t_0 and t_1).

For the step size control, set the initial step size $h = h_{max}$, and use the following strategy:

1. Take a step using the scheme and estimate τ
2. If $\tau \leq \delta h$ accept the step, otherwise reject the step
3. Assuming that $\tau(h) = Kh^5$, find a new step size h_{new} such that $\tau(h_{new}) = 0.5 \cdot \delta h_{new}$
4. Set $h \leftarrow \min(\min(\max(h_{new}, 0.1h), 4h), h_{max})$
5. If $h < h_{min}$, terminate with error message
6. If needed, adjust h so that the final step will end exactly at the final time $t = t_1$
7. Repeat until final time reached

Test the function using the script `erk1m1` on the course web page, only replacing the call to `ode45` with your function `rkck`. Set $abstol=1e-3$, $hmin=1e-6$, and $hmax=1.0$.

2. Write a MATLAB function with the syntax

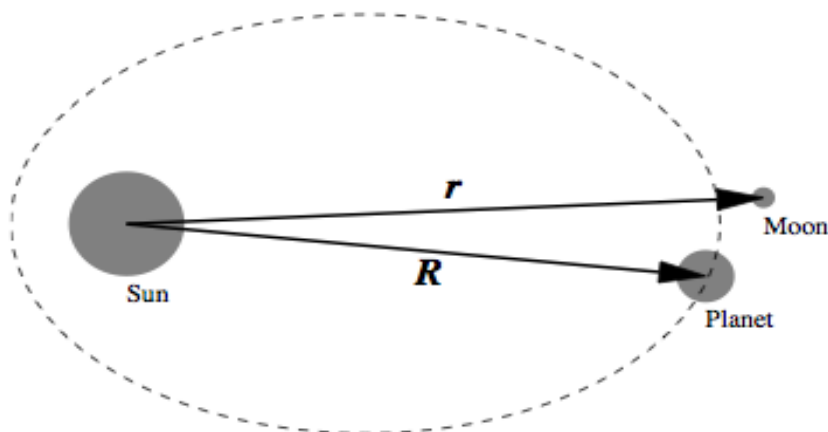
```
function [t,u]=orbit(m,T,u0)
```

which uses `rkck` from problem 1 to solve for the position up to time T of a planet $\mathbf{R} = (X, Y)$ of mass m_2 and a moon $\mathbf{r} = (x, y)$ of mass m_3 , starting from the initial conditions $u_0 = [X, Y, x, y, \dot{X}, \dot{Y}, \dot{x}, \dot{y}]^T$. Assume that only gravitational forces are present and that the sun of mass m_1 is fixed (see figure below). The (attractive) gravitational force between two bodies of mass m, \tilde{m} at $\mathbf{x}, \tilde{\mathbf{x}}$ is given by

$$\mathbf{F} = -m\tilde{m} \frac{\mathbf{x} - \tilde{\mathbf{x}}}{\|\mathbf{x} - \tilde{\mathbf{x}}\|^3}$$

and the motion $\mathbf{x}(t)$ of a body of mass m is given by Newton's second law, $m\ddot{\mathbf{x}} =$ the total force acting on the body. Use `abstol=1e-6`, `hmin=1e-6`, and `hmax=1.0`. Test the function using the commands

```
m=[1,1/100,1/8000]; T=30; u0=[5;0;5.1;0;0;.2;0;.3];
[t,u]=orbit(m,T,u0);
figure(1),plot(u(1,:),u(2,:),u(3,:),u(4:)),axis equal
figure(2),semilogy(t(1:end-1),diff(t))
```



3. a) Implement a MATLAB function

```
function c=mkfdstencil(x,xbar,k)
```

which computes the coefficients c_i for a finite difference approximation $u^{(k)}(\bar{x}) \approx \sum_i c_i u(x_i)$.

- b) Implement a MATLAB function

```
function u=bvp1(x,f,sigma,beta)
```

which solves the boundary value problem

$$u''(x) = f(x), \quad u'(a) = \sigma, \quad u(b) = \beta \quad (2)$$

using 3-point finite differences on a nonuniform grid. The inputs are the grid points \mathbf{x} , the right-hand side function $\mathbf{f}(\mathbf{x})$, and the boundary values. The output \mathbf{u} is the solution vector. Test the function using the commands

```

x=linspace(0,1,100).^2;
f=@(x) sin(2*pi*x).*exp(x);
u=bvp1(x,f,0,0);
plot(x,u)

```

c) Implement a MATLAB function

```
function [e1,e2,slope1,slope2]=bvpconv(ns)
```

which solves the problem in **b** using $f(x) = e^x$, $\sigma = -5$, $\beta = 3$, $a = 0$, $b = 3$, and grids with n points for each values n in the vector **ns**. Use the following two point distributions:

1. $x=3*\text{linspace}(0,1,n).^2$;
2. $x=3*\text{sort}(\text{rand}(1,n))$; $x(1)=0$; $x(n)=3$;

The outputs **e1,e2** should be the infinity norms of the errors for the two grid types (compare with the true solution), and **slope1,slope2** should be the estimated convergence rates (the slopes of the error versus $1/n$ in a log-log graph). Test the function using the commands

```

ns=round(logspace(1,3,50));
[e1,e2,s1,s2]=bvpconv(ns);
loglog(1./ns,e1,1./ns,e2)
[s1,s2]

```