

# Long-time behavior in geometric flows

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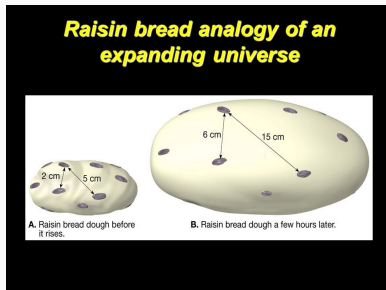
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# The big question

When does a geometric flow make a space more homogenous?

Today: Two ways to evolve a three dimensional space.

1. If a three dimensional space admits a locally homogeneous structure, can we find it with the Ricci flow?
2. For a compact universe with no matter, will gravitational dynamics make it more homogeneous?



# Outline of the talk

1. Homogeneous spaces and the geometrization conjecture
2. The geometrization conjecture and Ricci flow
3. Finiteness of the number of surgeries
4. Long-time behavior of Ricci flow
5. The Einstein flow

Homogeneous spaces and the geometrization conjecture

Geometrization conjecture and Ricci flow

Finiteness of the number of surgeries

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# Topology and geometric flows in three dimensions

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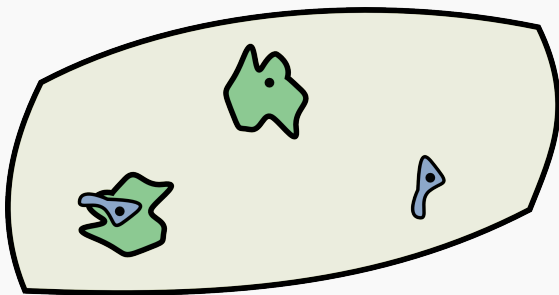
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First, how do we understand three dimensional spaces?

In terms of **homogeneous spaces**.

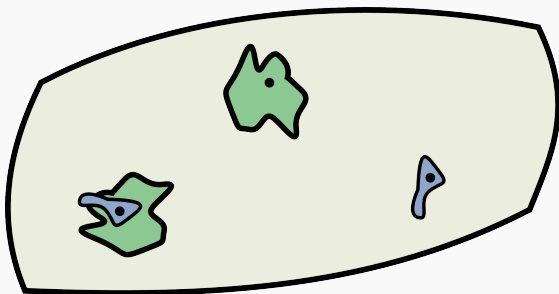
# Locally homogeneous metric spaces

A metric space  $X$  is *locally homogeneous* if all  $x, y \in X$ , there are neighbourhoods  $U$  and  $V$  of  $x$  and  $y$  and an isometric isomorphism  $(U, x) \rightarrow (V, y)$ .



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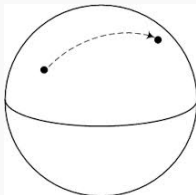
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The metric space  $X$  is *globally homogeneous* if for all  $x, y \in X$ , there is an isometric isomorphism  $\phi : X \rightarrow X$  that  $\phi(x) = y$ .

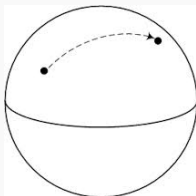
# Locally homogeneous Riemannian manifolds

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## Theorem

*(Singer 1960) If  $M$  is a complete, simply connected Riemannian manifold which is locally homogeneous, then  $M$  is globally homogeneous.*

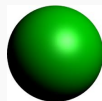
So passing to the universal cover turns “locally homogeneous” into “globally homogeneous”.

# Two-dimensional geometries

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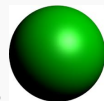
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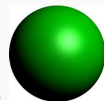
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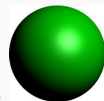


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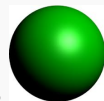
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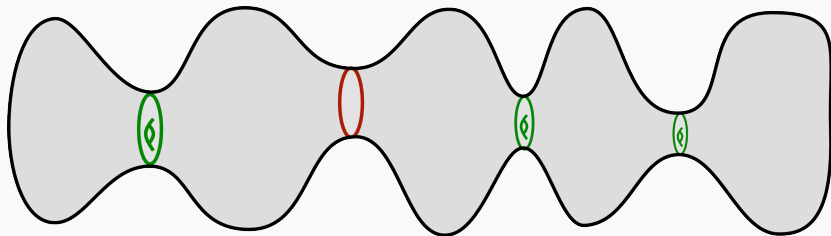
**Warning :** Unlike in two dimensions, not every compact three-dimensional manifold admits a geometric structure, i.e. admits a locally homogeneous Riemannian metric.

# Geometrization conjecture

If  $M$  is a compact orientable 3-manifold then there is a way to split  $M$  into canonical pieces, using certain embedded 2-spheres and 2-tori. (The collection of 2-spheres and 2-tori could be empty.)

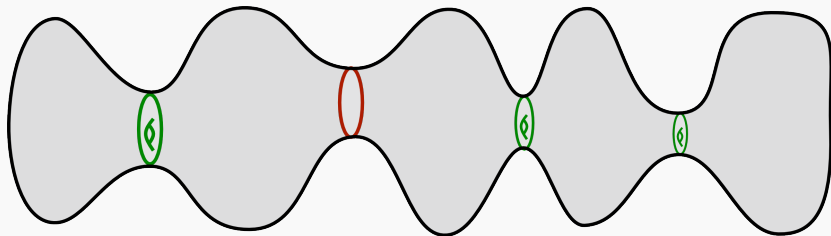
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Conjecture (Thurston, 1982)

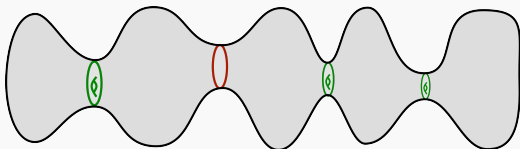
*The ensuing pieces have geometric structures, i.e. admit locally homogeneous metrics*

# Geometric decomposition

Cut along the 2-spheres and cap off the resulting pieces with 3-balls.

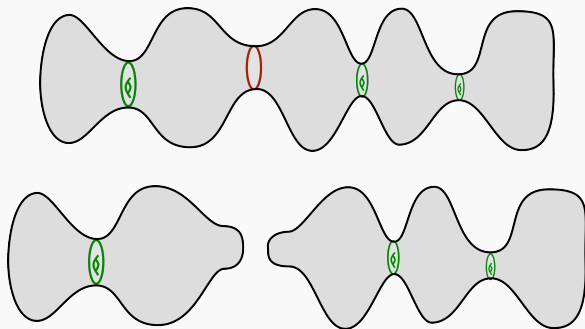
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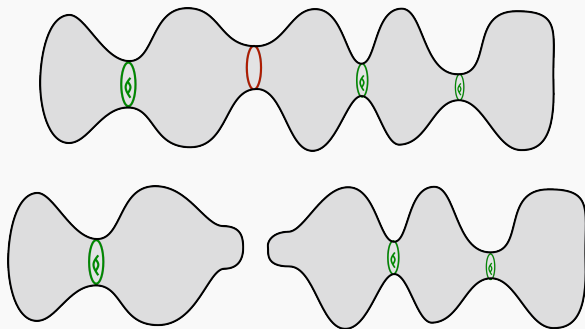
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Cut along the 2-tori. The interiors of the ensuing pieces should admit complete locally homogeneous metrics.



Homogeneous spaces and the geometrization conjecture

**Geometrization conjecture and Ricci flow**

Finiteness of the number of surgeries

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## Hamilton's Ricci flow equation

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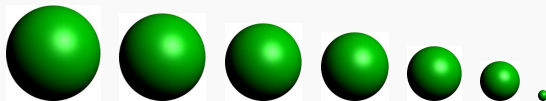
*Maybe* the Ricci flow will evolve an initial Riemannian metric into something homogeneous.

# Surfaces

For the Ricci flow on a compact surface, after rescaling the metric approaches a locally homogeneous metric.

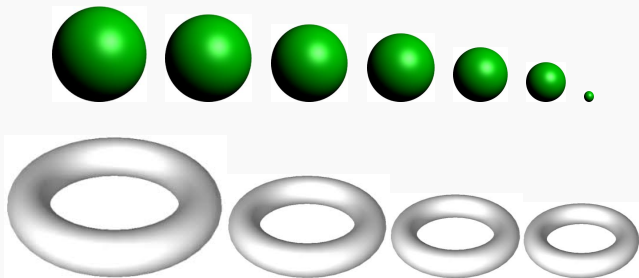
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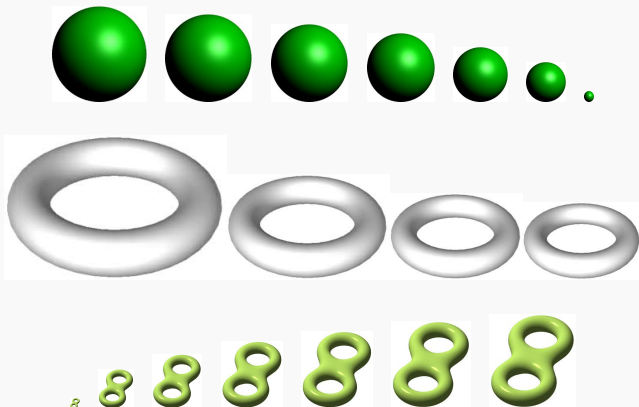
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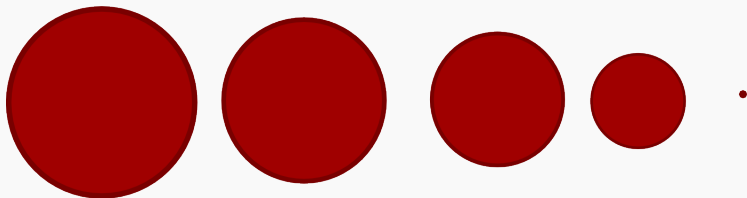
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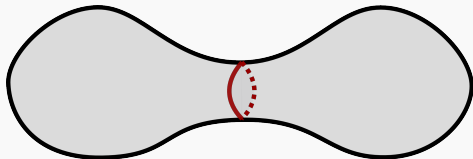
# Singularities in 3D Ricci flow

Some components may disappear, e.g. a round shrinking 3-sphere.



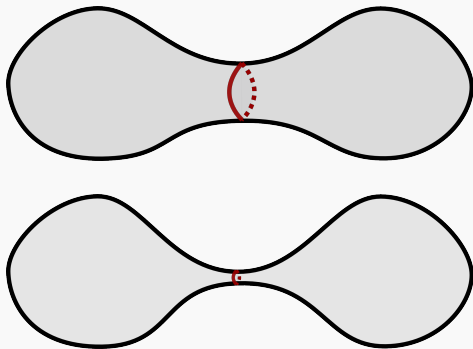
# Neckpinch

A 2-sphere pinches off. (Drawn one dimension down.)



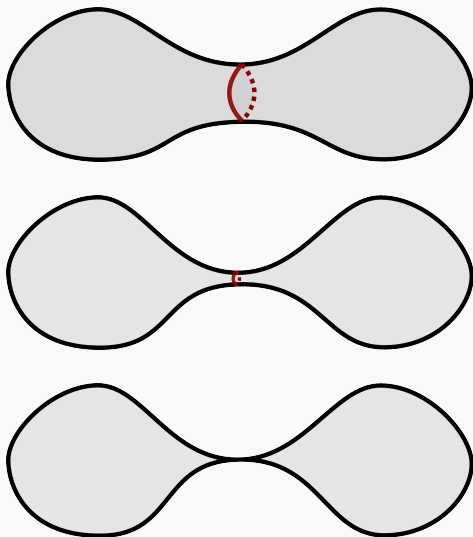
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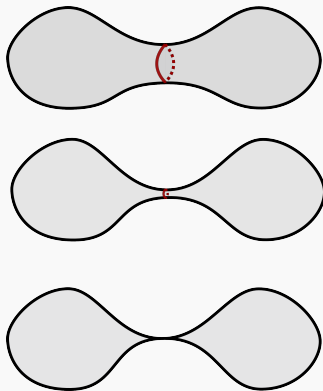


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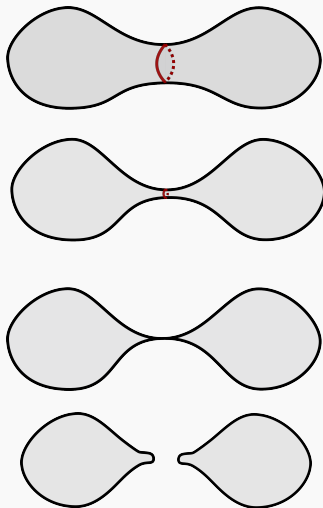
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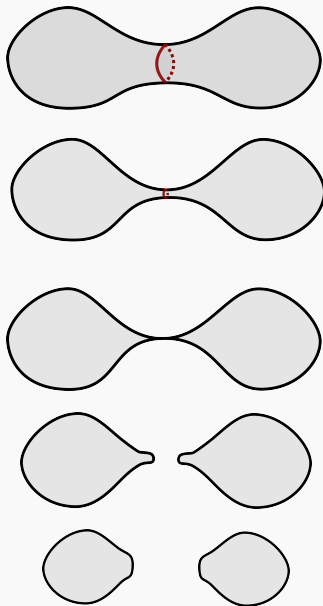
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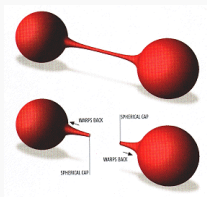


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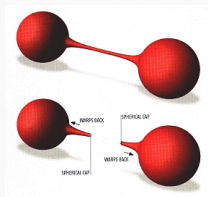




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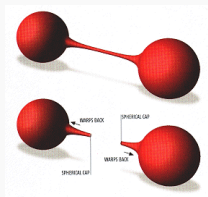


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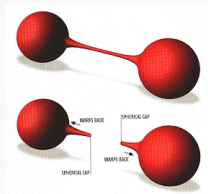
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Remark : the surgeries are done on 2-spheres, not 2-tori.

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**Step 3 :** Show that after the singularities are over, as time evolves, the locally homogeneous pieces in the Thurston decomposition asymptotically appear.

(Relevant geometries :  $\mathbb{R}^3$ ,  $H^3$ ,  $H^2 \times \mathbb{R}$ ,  $\widetilde{SL(2, \mathbb{R})}$ , Sol, Nil.)

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**From Perelman's first Ricci flow paper :** *Moreover, it can be shown ... that the solution is smooth (if nonempty) from some finite time on.*

**From Perelman's second Ricci flow paper :** *This is a technical paper, which is a continuation of [1]. Here we verify most of the assertions, made in [1, §13]; the exceptions are ... the smoothness of the solution from some time on, which turned out to be unjustified, and, on the other hand, irrelevant for the other conclusions.*

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2. The thin part is a “graph manifold”. (This doesn’t use Ricci flow. Stated by Perelman, proofs by Shioya-Yamaguchi, Morgan-Tian, Bessières-Besson-Boileau-Maillot-Porti and Kleiner-L.)

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Graph manifolds were known to have a geometric decomposition. Along with knowledge of the topological effects of surgeries, this proved the geometrization conjecture.

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**Remark** : Answering these questions has no topological implication. We already know that the geometrization conjecture holds. Rather, they are *analytic* questions about the Ricci flow.

# Long-time behavior

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Geometrization conjecture and Ricci flow

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The statement is that if this parameter is small enough (which can always be achieved) then there is a finite number of surgeries.

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This is good because we know lots about metrics with bounded sectional curvature (Cheeger-Fukaya-Gromov).

# Ingredients of the proof

Bamler's proof uses all of Perelman's work, and more. Some of the new ingredients :

1. Localizing Perelman's estimates and applying them to local covers of the manifold.
2. Use of minimal surfaces to control the geometry of the thin part.
3. Use of minimal embedded 2-complexes.

# Long-time behavior

Homogeneous spaces and the geometrization conjecture

Geometrization conjecture and Ricci flow

Finiteness of the number of surgeries

**Long-time behavior of Ricci flow**

Einstein flow

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The only case that we completely understand is when  $M$  admits *some* hyperbolic metric. Then from Perelman's work, for *any* initial metric on  $M$ , as  $t \rightarrow \infty$  the rescaled Riemannian metric  $\widehat{g}(t)$  approaches the metric on  $M$  of constant sectional curvature  $-\frac{1}{4}$ .



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**Question :** if  $M$  doesn't admit a hyperbolic metric, what are the candidate geometries for the long-time behavior?

# Quasistatic solutions

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Fact : On a compact 3-manifold, any self-similar solution has constant sectional curvature.

**Apparent paradox** : What happens to the Ricci flow if our 3-manifold doesn't admit a constant curvature metric?

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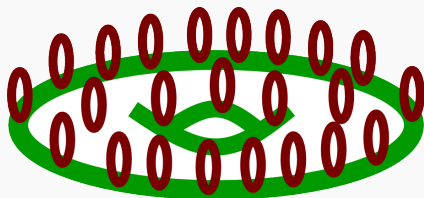
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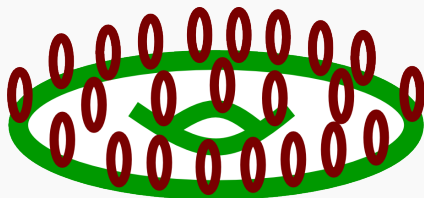
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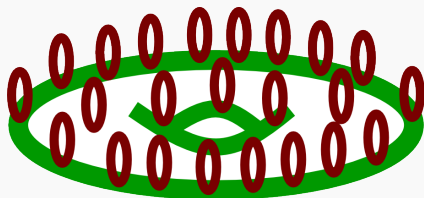


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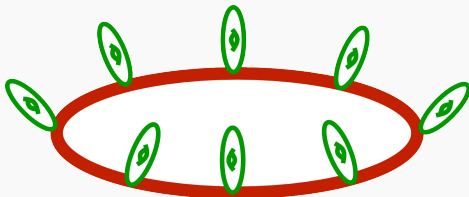
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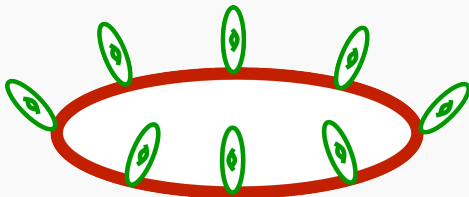
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With the rescaled metric  $\hat{g}(t) = \frac{g(t)}{t}$ ,  $(M, \hat{g}(t))$  shrinks to a point.

$M$  fibers over a circle with 2-torus fibers. The monodromy is a hyperbolic element of  $SL(2, \mathbb{Z})$ .

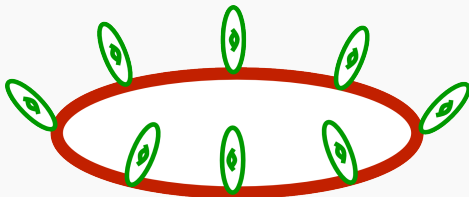


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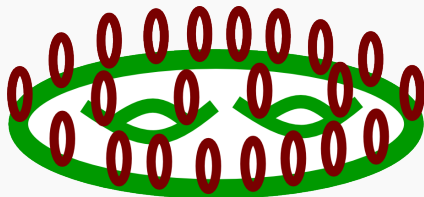
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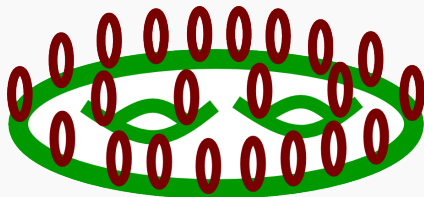
With the rescaled metric,  $(M, \hat{g}(t))$  approaches a circle.

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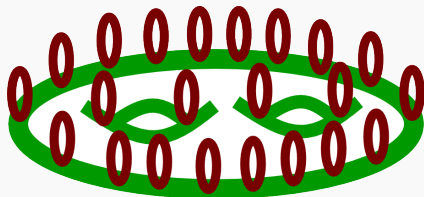


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With the rescaled metric,  $(M, \hat{g}(t))$  approaches the hyperbolic surface  $\Sigma$ . As the fibers shrink, the local geometry of the total space becomes more product-like.

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There is a common pattern, but to see it one must pass to the universal cover. That is, we are looking at the Ricci flow on a Thurston geometry of type  $\mathbb{R}^3$ ,  $H^3$ ,  $H^2 \times \mathbb{R}$ , Sol, Nil or  $\widetilde{SL_2(\mathbb{R})}$ .

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# The limiting solitons

<u>Thurston type</u>	<u>Expanding soliton</u>
$H^3$	$4 t g_{H^3}$
$H^2 \times \mathbb{R}$ or $\widetilde{\text{SL}(2, \mathbb{R})}$	$2 t g_{H^2} + g_{\mathbb{R}}$
Sol	$e^{-2z} dx^2 + e^{2z} dy^2 + 4 t dz^2$
Nil	$\frac{1}{3t^{\frac{1}{3}}} \left( dx + \frac{1}{2} y dz - \frac{1}{2} z dy \right)^2 + t^{\frac{1}{3}} (dy^2 + dz^2)$
$\mathbb{R}^3$	$g_{\mathbb{R}^3}$



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*(L. 2010) Suppose that  $(M, g(t))$  is a Ricci flow on a compact three-dimensional manifold, that exists for  $t \in [0, \infty)$ . Suppose that the sectional curvatures are  $O(t^{-1})$  in magnitude, and the diameter is  $O(\sqrt{t})$ . Then the pullback of the Ricci flow to  $\tilde{M}$  approaches one of the homogeneous expanding solitons.*

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## Conjecture

*For a long-time 3D Ricci flow, the diameter is  $O(\sqrt{t})$  if and only if  $M$  admits a locally homogeneous metric.*

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4. Local stability results for certain expanding Ricci solitons (due to Dan Knopf).



# A more refined result

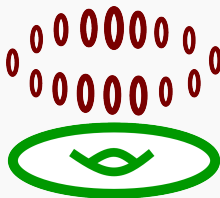
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**Question :** is this true for *all* initial metrics on  $T^3$ ?

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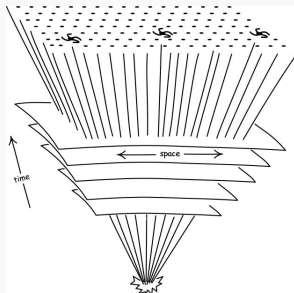
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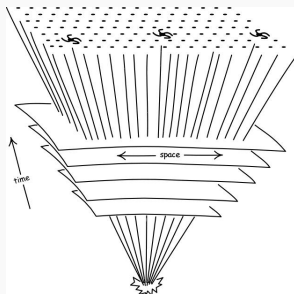
**Einstein flow**

# The setup



I'm interested in expanding vacuum spacetimes. What is the future behavior?

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The spacetime is diffeomorphic to  $(0, \infty) \times X$ , where  $X$  is a compact three-dimensional manifold.

# Einstein equations

The spacetime has a Lorentzian metric  $g$ . The Einstein equation of general relativity is

$$R_{\alpha\beta} - \frac{1}{2}Rg_{\alpha\beta} + \Lambda g_{\alpha\beta} = \frac{8\pi G}{c^4}T_{\alpha\beta}.$$

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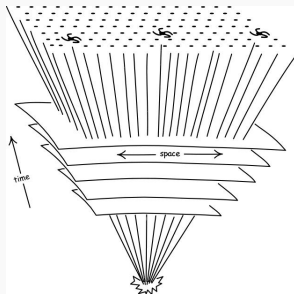
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Then the vacuum Einstein equation becomes

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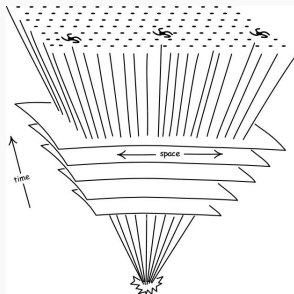
i.e. the Lorentzian metric  $g$  is Ricci-flat.

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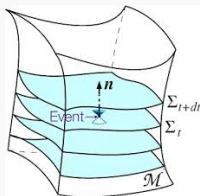


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There has been lots of work on this, mostly under some symmetry assumptions for the spatial slices (e.g. locally homogeneous or  $T^2$ -symmetry). Are there more general results?

# What is time?

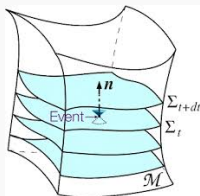
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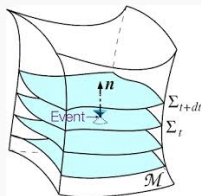
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Let's *assume* that along any given hypersurface, the expansion factor is constant. This defines a constant mean curvature (CMC) foliation.

# Einstein flow

Using the foliation, the metric takes the form

$$g = -L^2 dt^2 + h(t),$$

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where  $L = L(t)$  is a function on  $X$  and  $h(t)$  is a Riemannian metric on  $X$ . The Ricci-flat condition on  $g$  becomes

$$\frac{\partial h_{ij}}{\partial t} = -2LK_{ij} \quad (3)$$

and

$$\frac{\partial K_{ij}}{\partial t} = LHK_{ij} - 2L \sum_{k,l} h^{kl} K_{ik} K_{lj} - L_{;ij} + LR_{ij}, \quad (4)$$

along with certain time-independent “constraint” equations. Here the mean curvature  $H = \sum_{i,j} h^{ij} K_{ij}$  is spatially constant.

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The analogous statement in Ricci flow is that  $t^{-\frac{3}{2}} \text{vol}(X, h(t))$  is monotonically nonincreasing.



# Self-similar solutions

A Lorentzian metric  $g$  is *self-similar* if there's a one-parameter group of diffeomorphisms  $\{\phi_s\}$  so that  $\phi_s^*g = e^{cs}g$ , for some  $c \in \mathbb{R}$ .

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This is the analog of an (expanding) Ricci soliton.

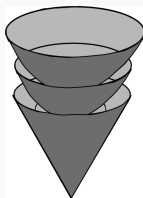
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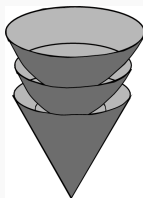
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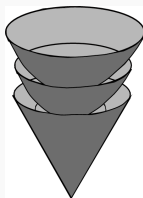


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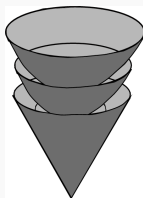


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2. The **Bianchi-III flat spacetime** is  $\mathbb{R}$  times the interior of a forward lightcone in  $\mathbb{R}^{2,1}$ .



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4. The **Kasner spacetimes** live on  $(0, \infty) \times \mathbb{R}^3$ , with metric

$$g = -dt^2 + t^{2p_1} dx^2 + t^{2p_2} dy^2 + t^{2p_3} dz^2.$$

Here

$$p_1 + p_2 + p_3 = p_1^2 + p_2^2 + p_3^2 = 1.$$

# Convergence result

The scale-invariant curvature condition is that  $\| \text{Rm}_g \| = O(t^{-2})$  as  $t \rightarrow \infty$ . (This is the analog of a type-III solution in Ricci flow.)

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*(L. 2018) Suppose that  $(h(t), K(t), L(t))$  is an expanding CMC Einstein flow on a compact aspherical three dimensional manifold  $X$ . Suppose that the curvature is  $O(t^{-2})$  in magnitude, and the diameter of  $(X, h(t))$  is  $O(t)$ .*

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(If there is a lower volume bound  $\text{vol}(h(t)) \geq \text{const} \cdot t^3$  then the model space is the Milne spacetime. This case is due to Mike Anderson.)

# Type-II solutions

Unlike in Ricci flow, there are expanding CMC Einstein flows that do *not* satisfy the scale-invariant curvature condition  $\| \text{Rm}_g \| = O(t^{-2})$ . (Homogeneous examples are due to Hans Ringström.)

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## Theorem

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*It turns out to be **flat**.*



# An apparent paradox

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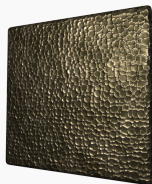
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In effect, there are increasing curvature fluctuations that average out the curvature to zero. The rescaled metrics *do* converge to a flat metric in the  $C^{1,\alpha}$ -topology.



# Some questions

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2. By Hawking's singularity theorem, if we look backward in time, there is geodesic incompleteness, and often curvature blowup. (Big bang.)

Can one understand the geometric asymptotics as one approaches the singularity? (BKL conjectures.)